

# Some Applications of Sparse Modelling in Physical Measurements

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Shiro Ikeda, ISM

## @ Today's Topics

@ X-ray diffraction imaging  
Phase retrieval

@ Astronomy data analysis  
Compton camera imaging

# Phase Retrieval

joint work with

H. Kono (JAEA)

# Background

- ① Fourier transform & phase
- ① 3D structure of biomolecule
- ① X-ray Free Electron Laser (XFEL)
- ① X-ray diffraction of single biomolecule

# Proposed method

- ① Phase retrieval
- ① Standard method (HIO)
- ① Proposed method (SPR)
- ① Numerical experiments



## ① 3D structure of biomolecule

### ① 1 dimensional structure

A chain of amino acid. (There are 20 possible amino acid)

ex) Lysozyme (a chain of 129 amino acid)

K V F G R C E . . . .

K: Lysine

V: Valine

F: Phenylalanine

G: Glycine

### ② 3 dimensional structure

The chain folds and forms a 3d structure

The reaction of the protein is determined by the 3d structure.

3d structure is important for medicine, biology, etc. . . .

## Protein 3D structure of Lysozyme

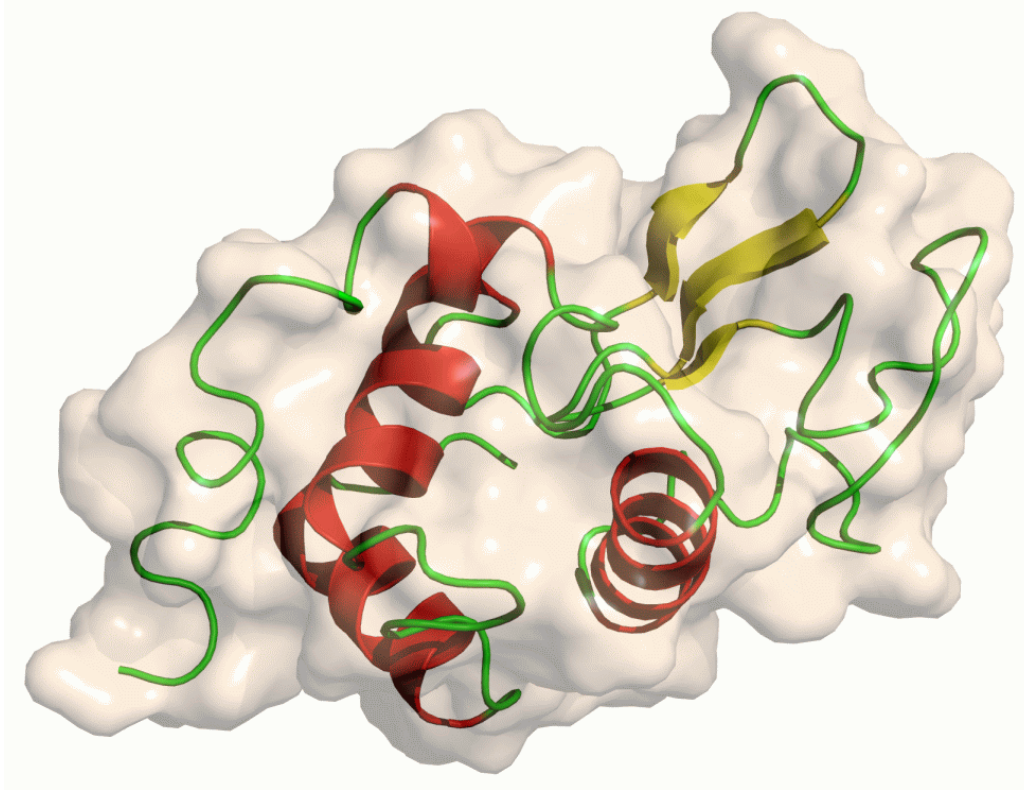


Figure: Lysozyme (wikipedia)

# Protein 3D structures

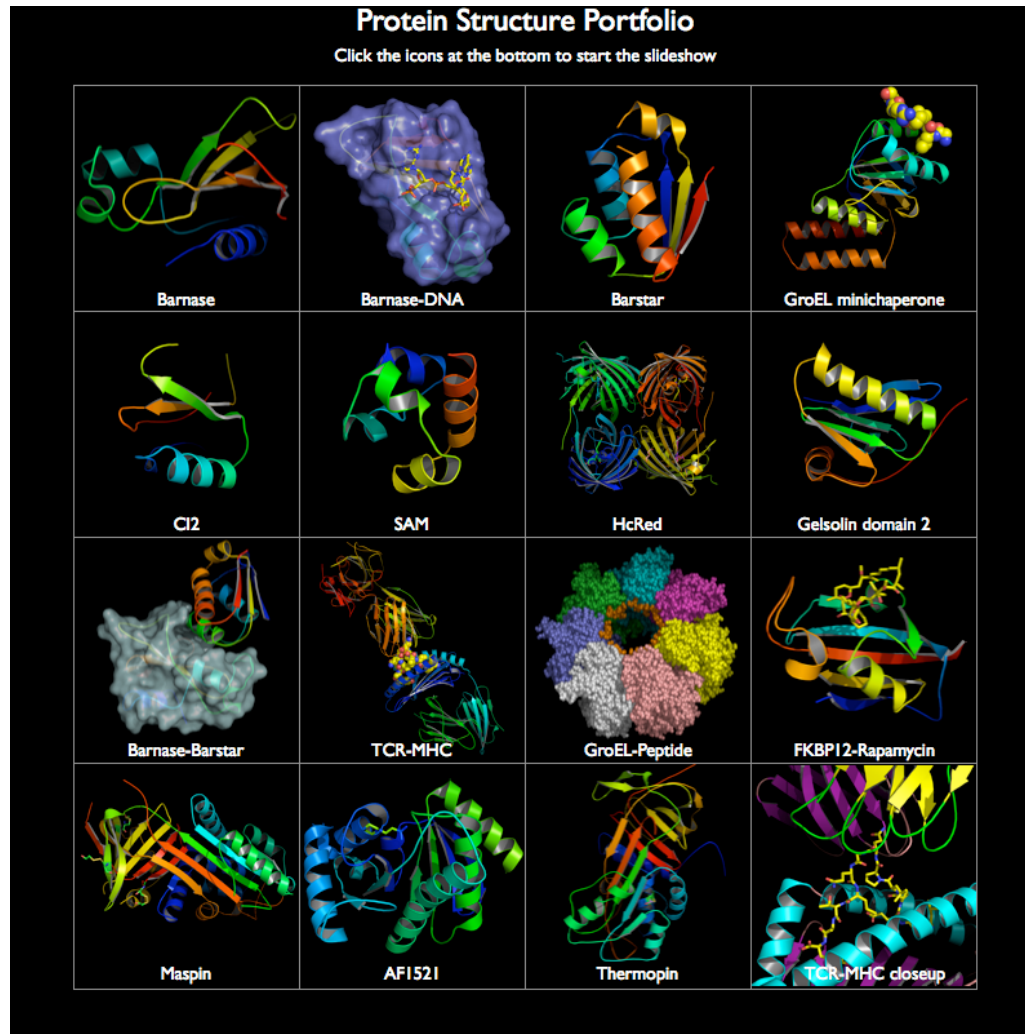


Figure: Protein structure gallery (Monash univ.)

## Conventional method

### ① Molecular Dynamics (MD)

② Simulate atomic dynamics with a super computer

③ Difficult for large molecules (few  $\mu\text{sec}$ )

### ① X-ray Crystallography (from 1950's)

Crystalization, X-ray Diffraction  $\rightarrow$  Analysis

② Main subject is the Crystalization

③ Many protein does not become a crystal

40% of biomolecule will not become a crystal.

## Crystallization

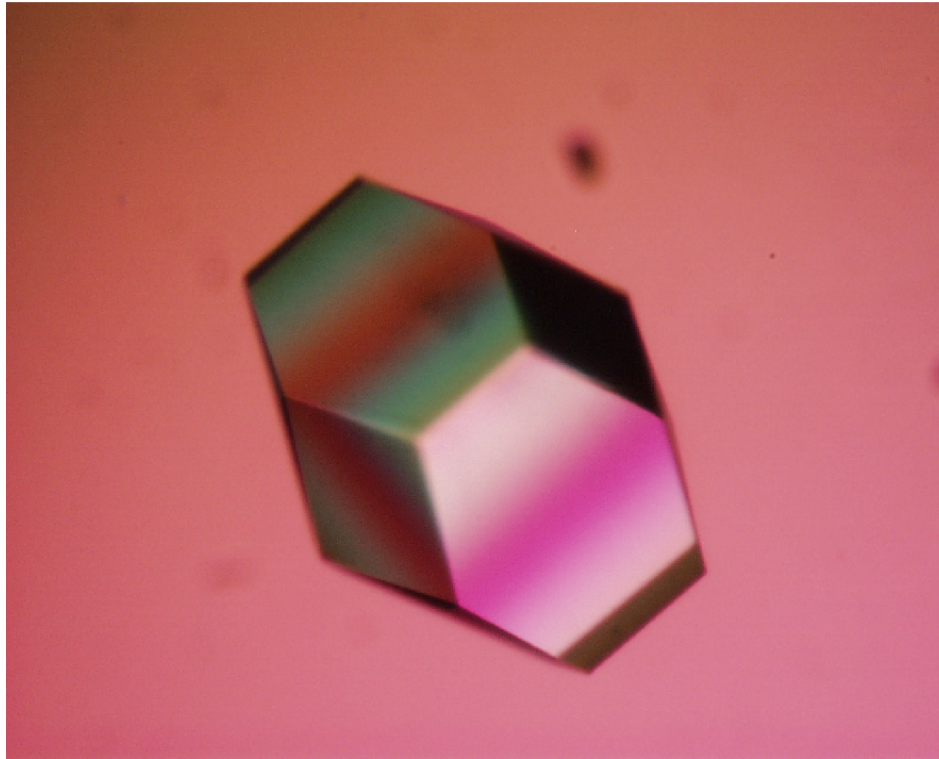
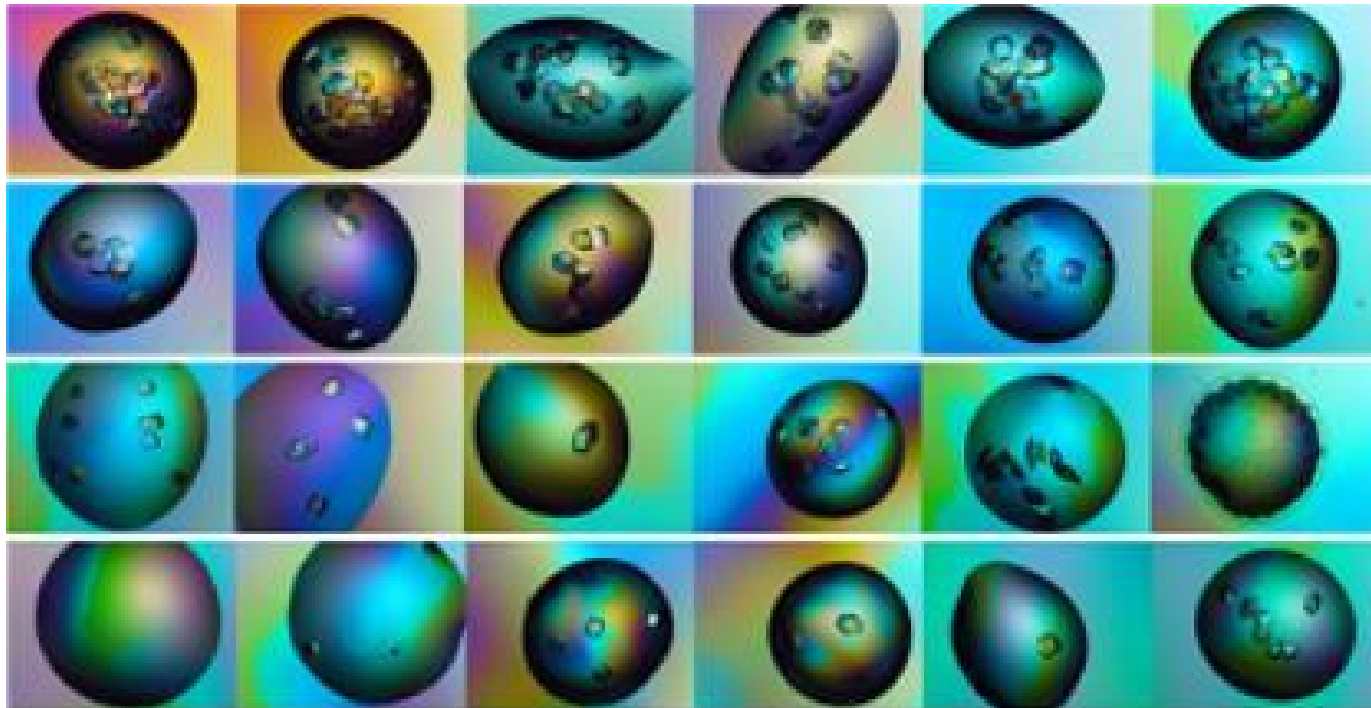


Figure: Lysozyme (wikipedia)

# Crystalization



Growing Protein Crystals

Figure: Crystalization.(Cornell Univ. CrySis)

## Crystalization and protein structure identification

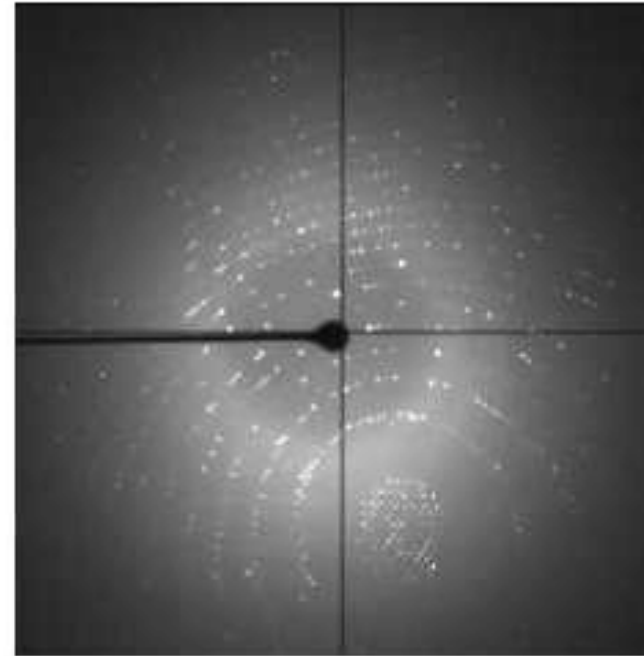
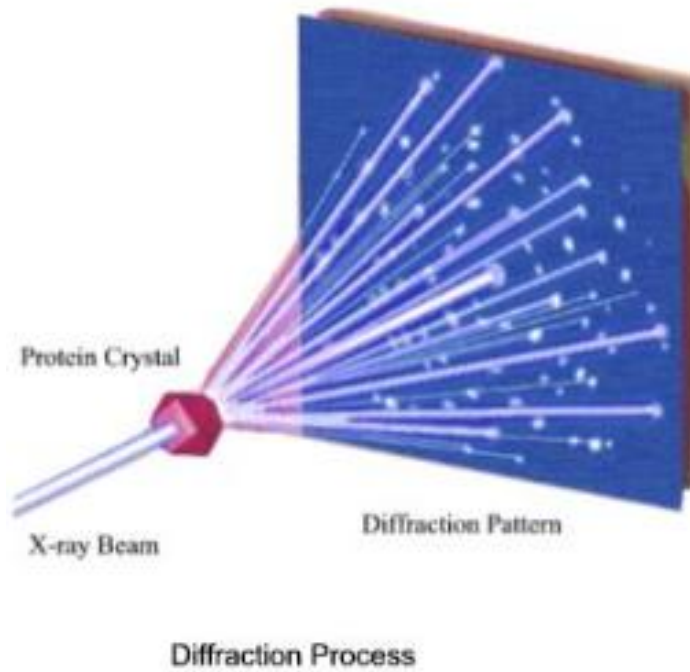


Figure: Crystalization and diffraction pattern.(Cornell Univ. CrySis)

# Crystallization and protein structure identification

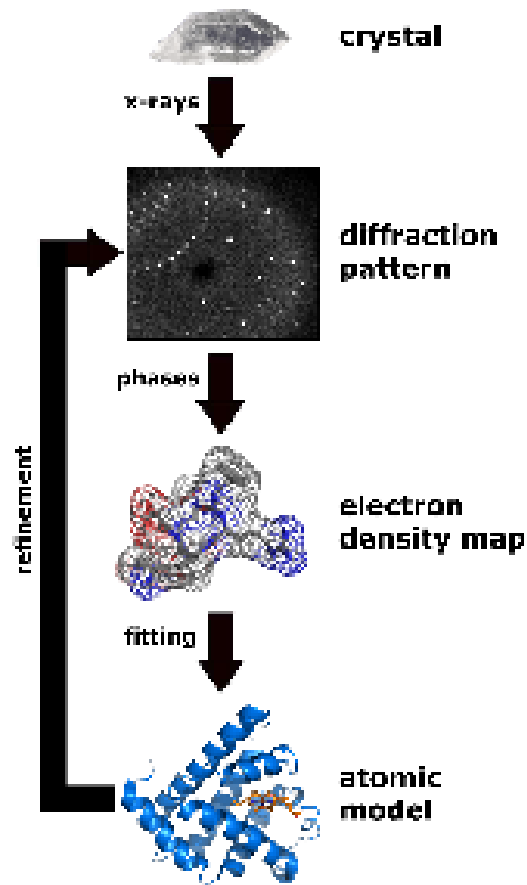


Figure: Crystallization and Structure recovery.(wikipedia)



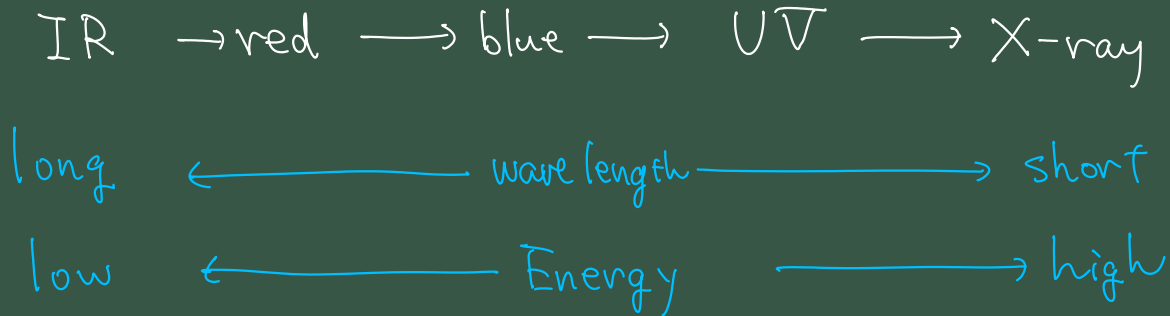
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## Q X-ray Free Electron Laser



gas laser, solid-state laser, and semiconductor laser, cannot emit high power X-ray laser.

## XFEL (X-ray free electron laser)



Figure: Spring8

## XFEL (X-ray free electron laser)



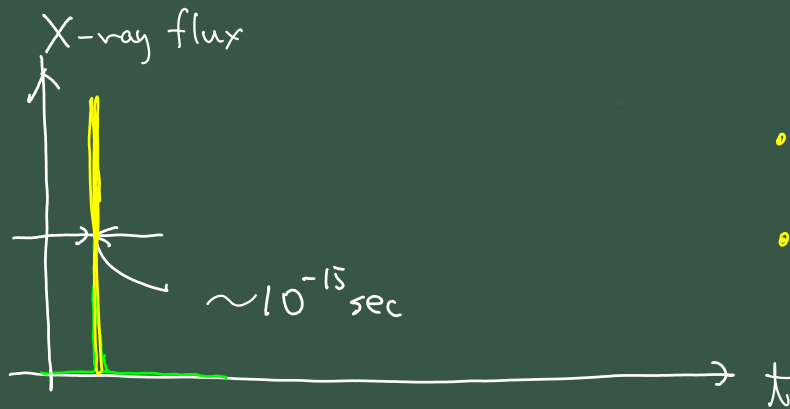
Figure: XFEL

# X-ray Free Electron Laser

① 4 projects in the world.

② Japan "lased" after USA.

(Cheaper, More Flux, Larger Energy)



- Short pulse
- Large flux X-ray

## XFEL (X-ray free electron laser)



Figure: XFEL in the world

# Background

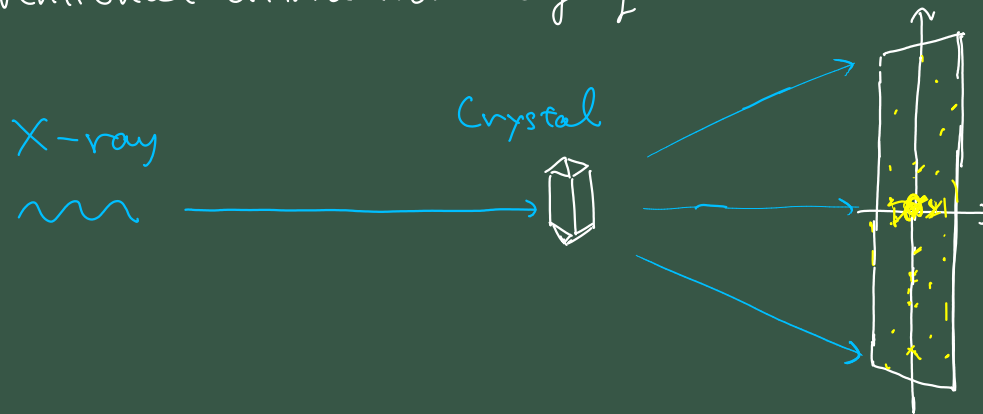
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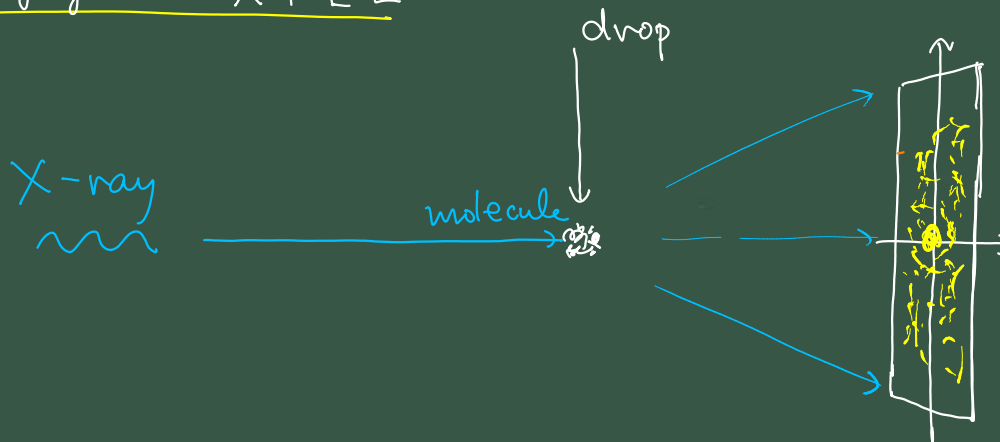
# ① Single molecule diffraction with XFEL

## Conventional diffraction imaging



- ① X-ray is not strong but repeat until we have good image

## Imaging with XFEL



- ① hit the X-ray with free falling molecule
- ② Experiment will start soon



# Protein structure identification with XFEL

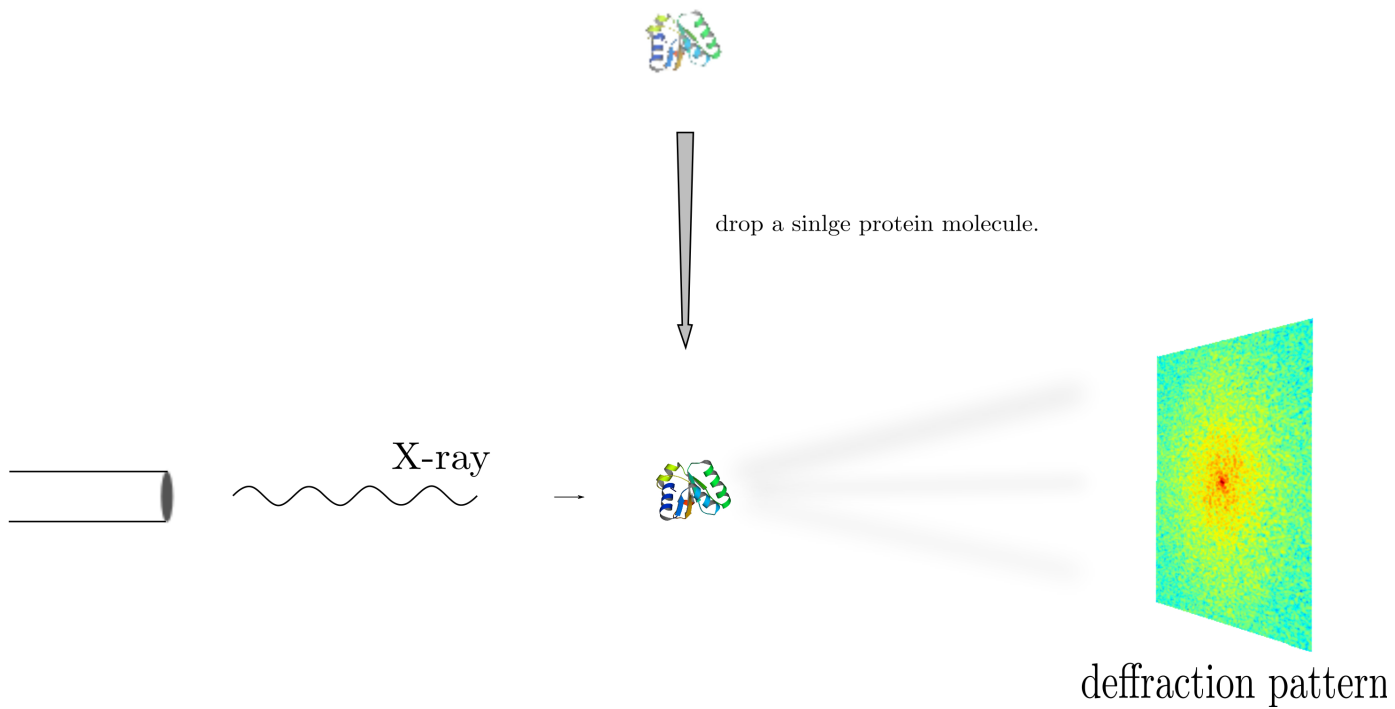


Figure: Protein diffraction pattern with XFEL.

# Problems

- ① Hit rate
- ② Molecule size and flux strength
- ③ Imaging from many angles. It is necessary to repeat measurement many times (10,000 ~ 100,000)
- ④ Relative angles between images.
- ⑤ Phase Retrieval

## Protein structure identification with XFEL

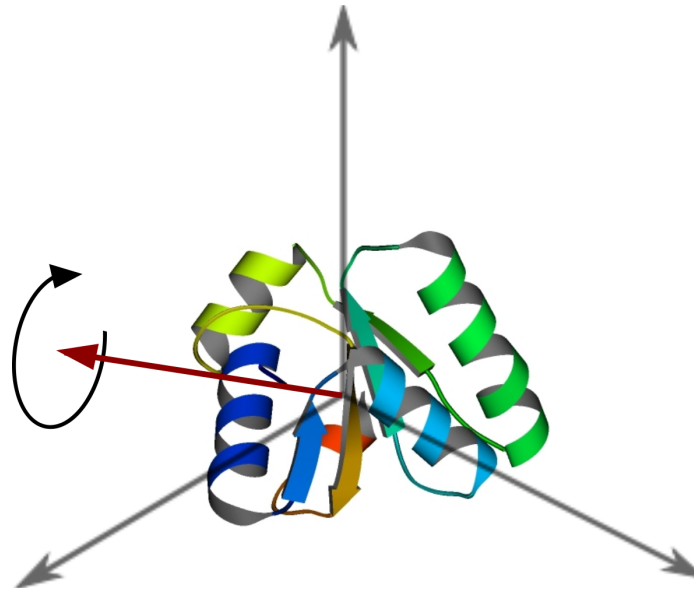


Figure: Angle retrieval problem.

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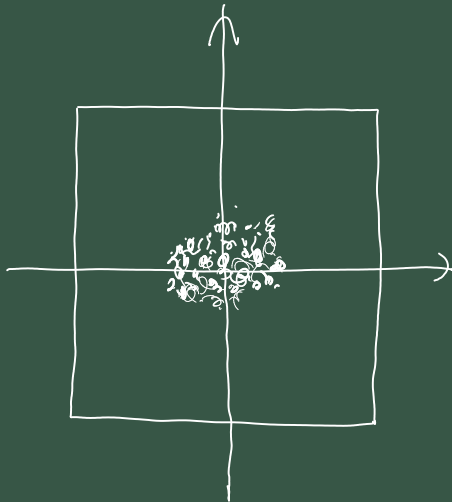
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# Q Phase Retrieval

Molecule  $\xrightarrow{\text{X-ray imaging}}$  Diffraction image

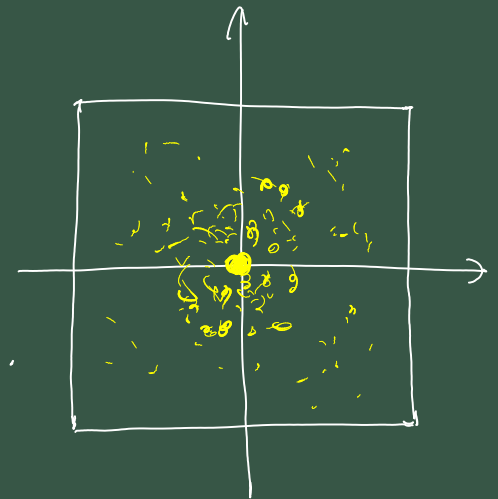
Electron density  $f_{xy}$   $\xrightarrow{\text{Fresnel diffraction}}$  Power Spectrum  
Fourier trans.  
 $F_i = \mathcal{F}(f)$   
 $|F_{uv}|^2$



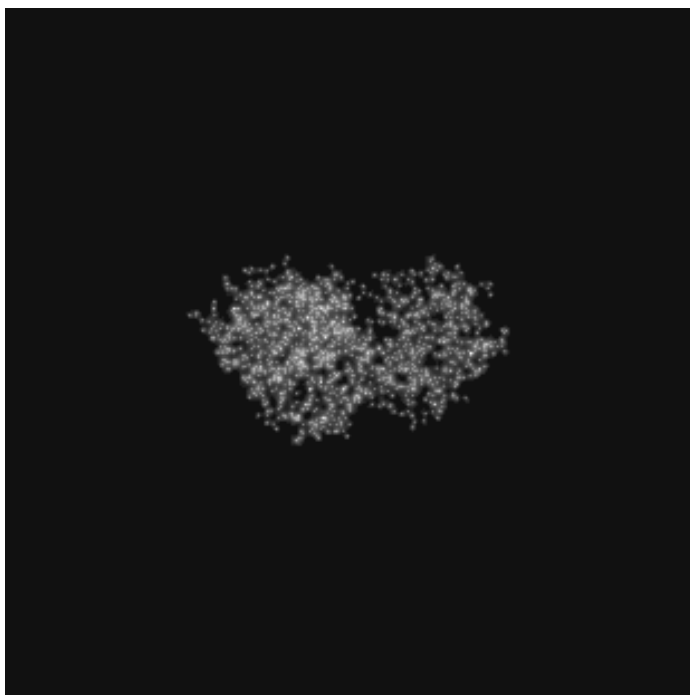
Inverse Fourier trans?



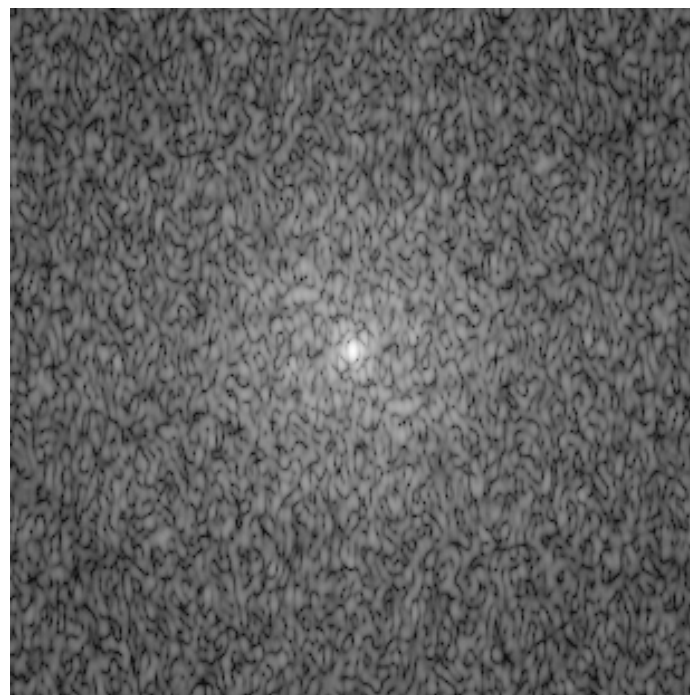
Phase is not known.



## Protein structure identification with XFEL



(a) Electron density.



(b) Ideal diffraction pattern.

Figure: Electron density and ideal diffraction pattern.

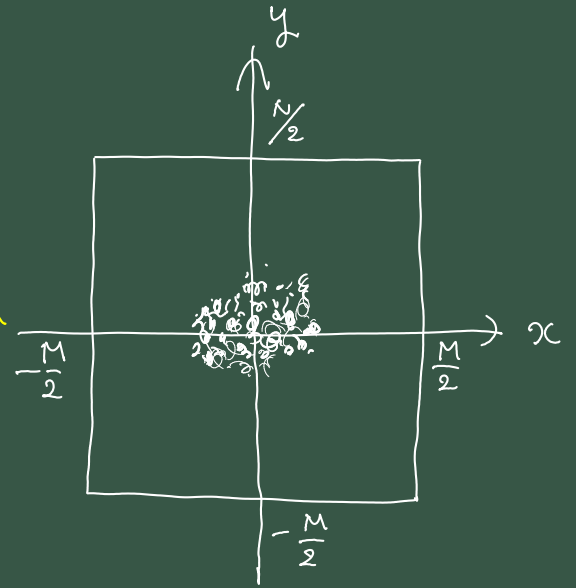
# Electron density & Diffraction image

Electron density (3 dim  $\rightarrow$  2 dim. projection)

$$f = \{f_{xy}\} \quad -\frac{M}{2} \leq x, y \leq \frac{M}{2}$$

$$f_{xy} \geq 0, f_{xy} \in \mathbb{R}$$

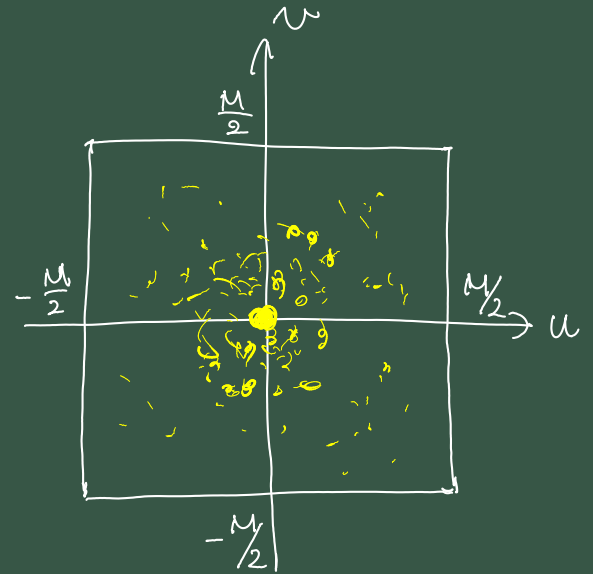
each pixel of electron density



## ④ Fourier Transform

$$F = \{F_{uv}\} \quad -\frac{M}{2} \leq u, v \leq \frac{M}{2}$$

$$F_{uv} = \mathcal{F}(f) = \frac{1}{M} \sum_{x,y} f_{xy} \exp\left(\frac{2\pi i (ux + vy)}{M}\right)$$



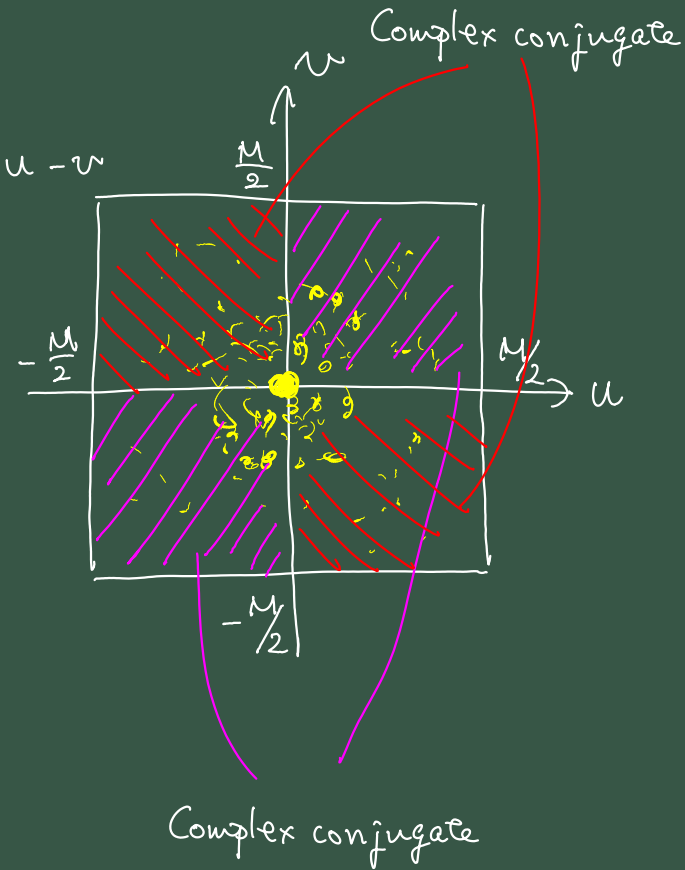
①  $f_{xy}$  is real number

$F_{uv}$  is the complex conjugate of  $F_{-u-v}$

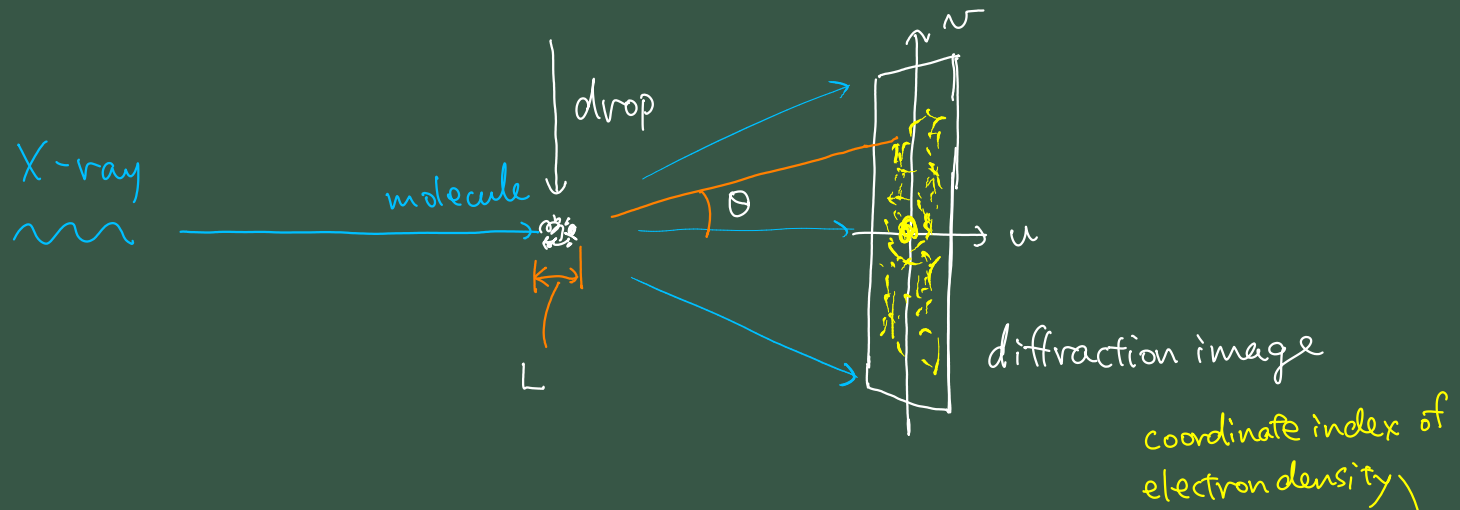
$$\therefore F_{uv} = \frac{1}{M} \sum_{x,y} f_{xy} \exp\left(\frac{2\pi i (ux+vy)}{M}\right)$$

$$\operatorname{Re}(F_{uv}) = \operatorname{Re}(F_{-u-v})$$

$$\operatorname{Im}(F_{uv}) = -\operatorname{Im}(F_{-u-v})$$







The relation between diffraction image  $S_{uv}$  and electron density  $f(x,y)$  is written as follows

pixel index of diffraction image  $\underbrace{S_{uv}}$  and coordinate index of electron density  $\underbrace{f(x,y)}$

$$S_{uv} = \alpha |F_{uv}|^2 \cos^3 \theta$$

$$F_{uv} = (\mathcal{F}(f))_{uv} \quad ; \text{ Fourier transform}$$

$$\alpha = I r_e^2 \left( \frac{\lambda}{2L} \right)^2, \quad I : \text{ X-ray flux (photons/pulse/mm}^2\text{)}$$

$\lambda$  : wave length

$r_e$  : electron radius

$\theta, L$  : see the figure

$S_{uv}$  corresponds to the power spectrum of  $f(x, y)$

$f \rightarrow S$  is easy ( $|(\mathcal{F}(f))_{uv}|^2 \cos^3 \theta$ )

$S \rightarrow f$  is difficult (Phase is not known)

$f$ :  $M \times M$  real matrix

$S$ :  $\frac{M \times M}{2}$  real matrix

) ill-posed problem

•  $S \rightarrow f$  : Phase retrieval method

J. Fienup, Applied Optics (1982)

HIO (Hybrid Input Output) method is widely used.

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# ① HIO method

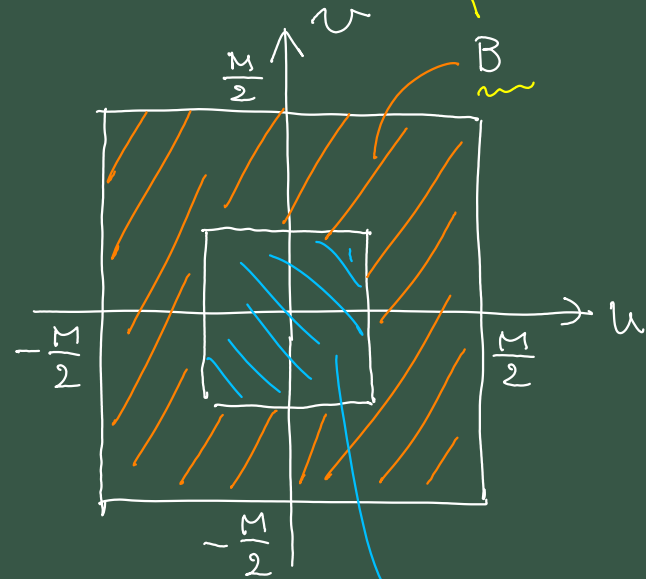
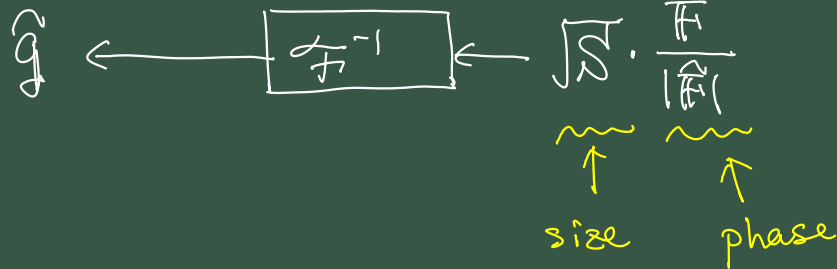
Let  $\hat{f}$  be the estimate of  $f$ .  $\hat{f}$  satisfies the followings

①  $\hat{f}_{xy} \geq 0, \therefore$  non-negative

②  $\hat{f}_{xy} = 0$  for  $(x, y) \in B, \therefore$  concentrate around center.



modify  $\hat{f}$  in order to satisfy ① and ②



$\hat{f}_{uv} \geq 0$   
around the center

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## ① From HIO to SPR (Sparse Phase Retrieval)

- HIO works well for noiseless data.
- For a large particle, diffraction patterns are strong and HIO works well.
- The target of XFEL is small particles and HIO will not work well.

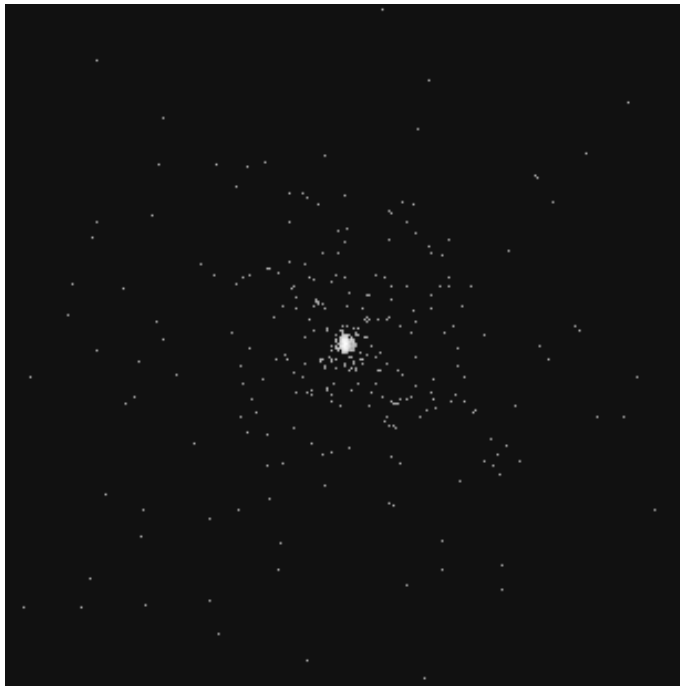


New method is needed.

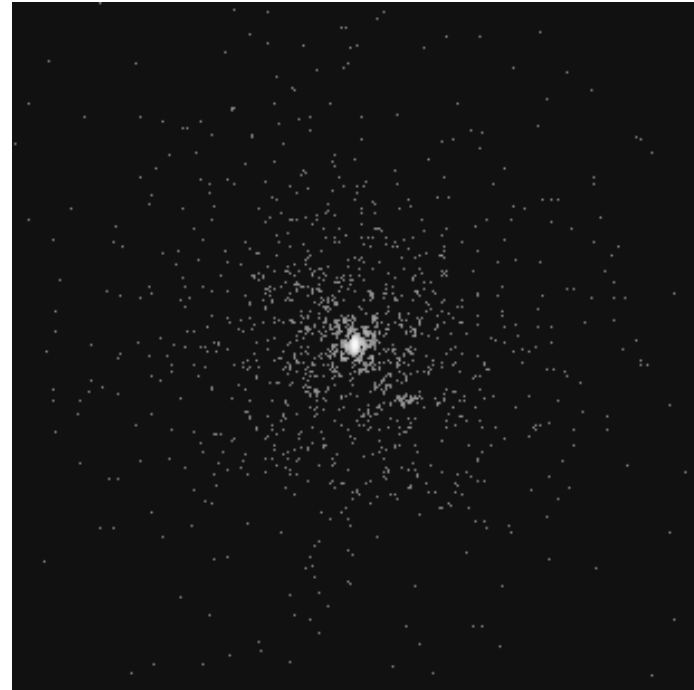
① Assume  $f$  is sparse

② Make noise model and use the sparse prior. Estimate base on Bayesian statistics.

## Protein structure identification with XFEL

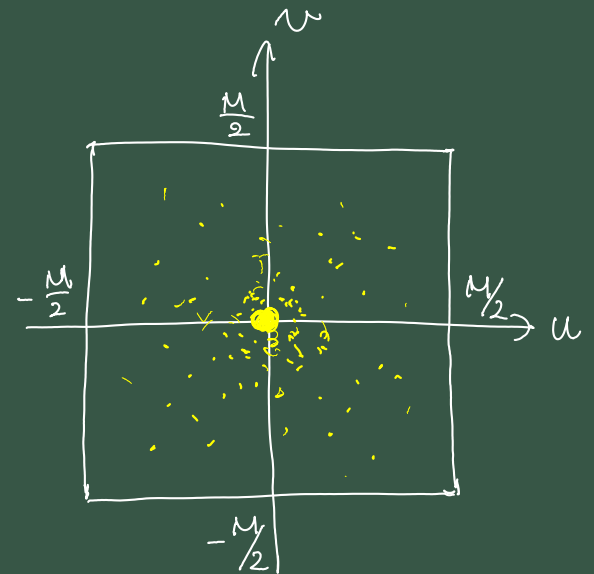
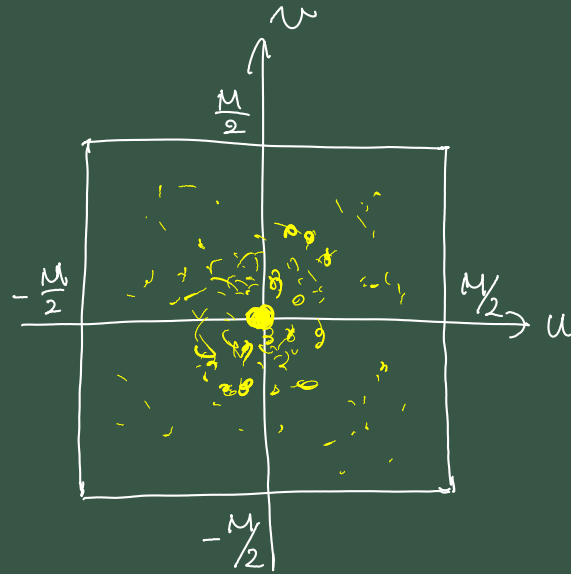
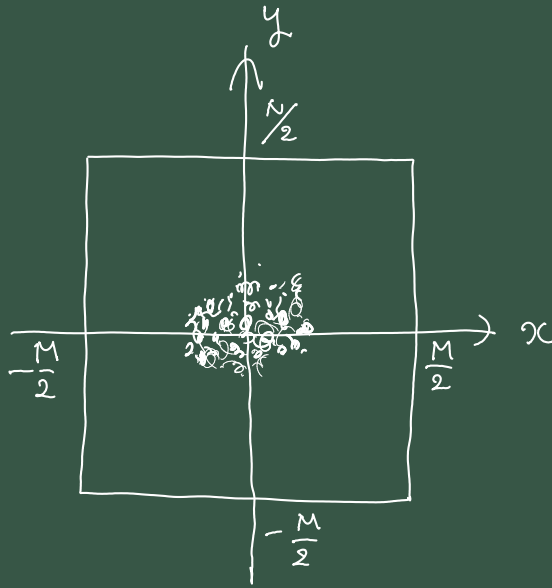


(a) X-ray fluxes  $1.0 \times 10^{21}$ (photons/pulse/mm<sup>2</sup>).



(b) X-ray fluxes  $5.0 \times 10^{21}$ (photons/pulse/mm<sup>2</sup>).

Figure: Diffraction patterns in experimental situation.



Electron density

$$f = \{f_{xy}\}$$

Ideal diffraction pattern

$$S = \{S_{uv}\}$$

$$S_{uv} = |F_{uv}|^2 \cos^2 \theta$$

$$F_{uv} = (\mathcal{F}(f))_{uv}$$

Measurement

$$N = \{N_{uv}\}$$

The number of photons  
at each pixel.

What is the relation between  $S_{uv}$  and  $N_{uv}$ ?



Measurement of each pixel is the number of photons which is an integer. Denoted as  $N_{uv}$

$N_{uv}$  : Nonnegative integer  $-\frac{M}{2} \leq u, v \leq \frac{M}{2}$ .

$N_{uv}$  follows a Poisson distribution whose expected value is  $S_{uv}$

$$p(N_{uv} | S_{uv}) = \frac{S_{uv}^{N_{uv}} \exp(-S_{uv})}{N_{uv}!}$$

$$p(\mathbb{N} | f) = \prod_{uv} \frac{(|F_{uv}|^2 c_{uv})^{N_{uv}} \exp(-|F_{uv}|^2 c_{uv})}{N_{uv}!}$$

(  $c_{uv} = \cos^3 \theta$  and  $S_{uv} = \alpha |F_{uv}|^2 c_{uv}$  where we set  $\alpha = 1$ )

$p(N | \theta)$  is modelled. How to estimate  $\theta$ .



Statistical Estimation Problem

① Maximum Likelihood Estimate (MLE)

Compute  $\theta$  which maximizes  $p(N | \theta)$

⇒ Ill posed problem  
(phase)

② Bayesian Statistics

Assume the prior of  $\theta$  as  $\pi(\theta)$

$\pi(\theta) \cdot p(N | \theta) = p(\theta, N)$  -- Joint dist

$$p(\theta | N) = \frac{p(\theta, N)}{\int p(\theta, N) d\theta}$$

posterior prob.

⇒ Maximize

MAP estimate

$$\underline{p(f|N)} \propto \underline{p(N|f)} \cdot \underline{\pi(f)} \quad (\underline{\text{Bayes Theorem}})$$

Posterior

Poisson model  
Likelihood

Sparsity  
Prior dist

MAP (Maximum a Posterior) estimate  $\hat{f}$  maximize  
above  $p(f|N)$

• maximize  $p(N|f) \cdot \pi(f) \Rightarrow$  maximize  $\log p(N|f) \pi(f)$

•  $\underline{\log p(N|f)} + \underline{\log \pi(f)}$

likelihood ↓

↓  
Prior

$$\underline{\sum_{uv} (N_{uv} \ln |F_{uv}|^2 - |F_{uv}|^2 C_{uv})} + \underline{\log \pi(f)}$$

How to define  $\pi(f)$  ?

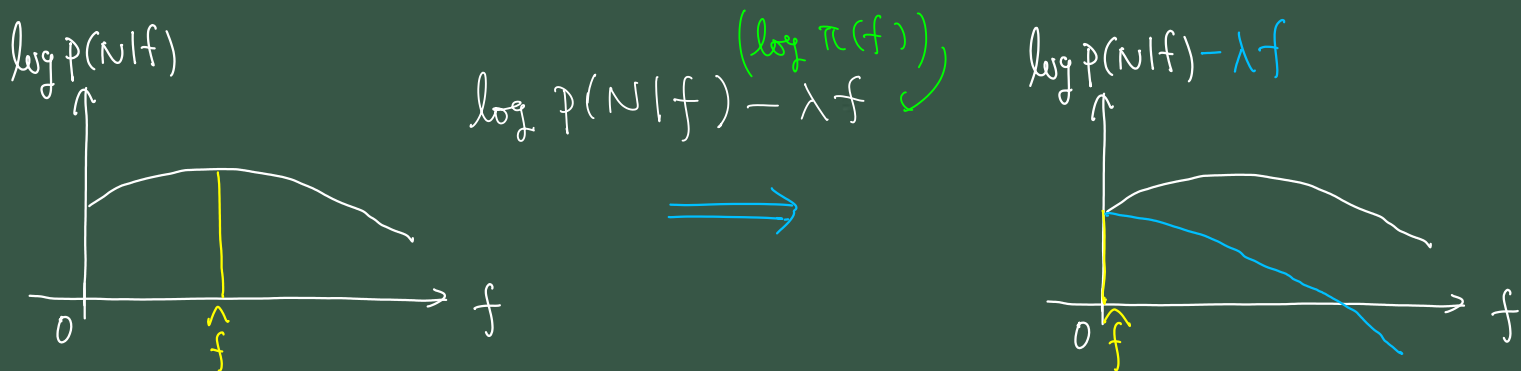
①  $\pi(f)$  reflects the prior knowledge of  $f$ .

②  $f$  has a lot of 0 components.  $\Rightarrow$  sparse

$$\pi(f) = \lambda \exp(-\lambda f) \quad (f \geq 0)$$

If we use above  $\pi(f)$ , a lot of  $f_{xy}$  become 0.

( LASSO : statistics 1996  
Compressed Sensing : Inform Th. 2000 ~ )



Define prior  $\pi(f)$  as follows

$$\pi(f) = \rho_{xy} \exp(-\rho_{xy} f_{xy}), \quad f_{xy} \geq 0.$$

$$\rho_{xy} = \mu \cdot w_{xy} = \underbrace{\mu}_{\substack{\mu \text{ is a hyper parameter}}} \cdot \frac{2}{N^2} (x^2 + y^2) \quad \rho_{xy} \rightarrow \text{large means } f_{xy} \text{ is more likely to be 0.}$$

Collect  $f$  related terms from  $\log p(f | N)$ ,

$$l(f | N) = \underbrace{\sum_{uv} (N_{uv} \ln |F_{uv}|^2 - |F_{uv}|^2 C_{uv})}_{\text{likelihood term}} - \underbrace{\mu \sum_{xy} w_{xy} f_{xy}}_{\text{prior term}}$$

$$\hat{f} = \arg \max_f l(f | N) \text{ is the problem.}$$

SPR

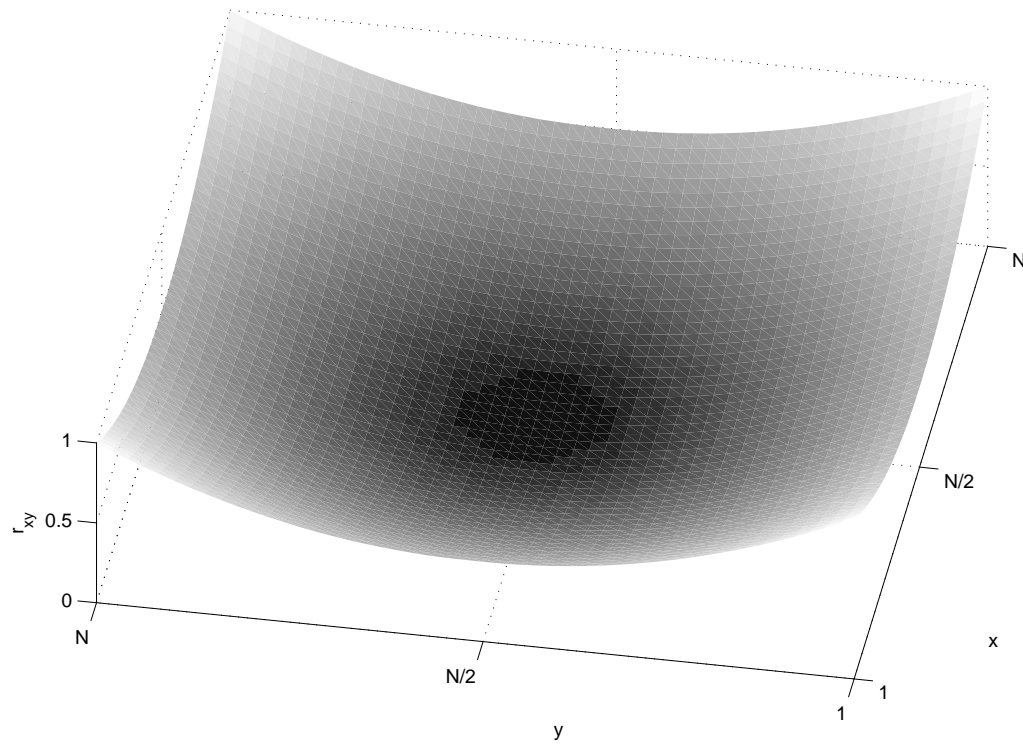


Figure: Sparsity prior  $w_{xy}$ .

$$Q(f|N) = \underbrace{\sum_{uv} (N_{uv} \ln |F_{uv}|^2 - |F_{uv}|^2 C_{uv})}_{\text{frequency domain}} - \underbrace{\mu \sum_{xy} W_{xy} f_{xy}}_{\text{real domain}}$$

mixed

Optimizing with a gradient-based method.

Phase retrieval is an ill-posed problem. It is necessary to use additional information

HIO method - - - - active region

SPR method - - - - sparsity

準備

$$\frac{\partial |F_{uv}|^2}{\partial f_{xy}} = \frac{2}{M} \operatorname{Re} \left( F_{uv} \exp \left( -\frac{2\pi i (ux+vy)}{M} \right) \right)$$

$$\text{Inverse Fourier trans. } \mathcal{F}^{-1}(F) = \frac{1}{M} \sum_{uv} F_{uv} \exp \left( -\frac{2\pi i (ux+vy)}{M} \right)$$

$$\frac{\partial \ell(f|N)}{\partial f_{xy}} = 2 \operatorname{Re} \left( \mathcal{F}^{-1}(g(F; N)) \right)_{xy} - \mu w_{xy}$$

$$, g_{uv}(F; N) = \left( \frac{Nuv}{|F_{uv}|^2} - C_{uv} \right) F_{uv}$$

$$f_{xy}^{t+1} = \max \left( 0, f_{xy}^t + \eta_t \frac{\partial \ell(f^t | N)}{\partial f_{xy}^t} \right)$$

Each  $f^t \rightarrow f^{t+1}$  needs a Fourier & inverse Fourier trans.

$\eta_t$  controls the step size. A simple line search speeds up the convergence.



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## Q Numerical experiments.

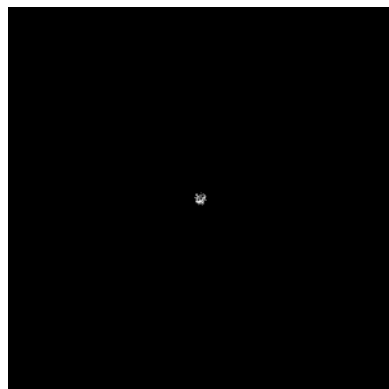
When  $\mu$  is large, many  $f_{xy}$  become 0 and the optimization is easy.

Starting with a large  $\mu$ , shrink it. For each  $\mu$ , compute  $\hat{f}$  and the following error( $\mu$ ). the  $\mu$  which minimizes the error is the optimal  $\mu$ .

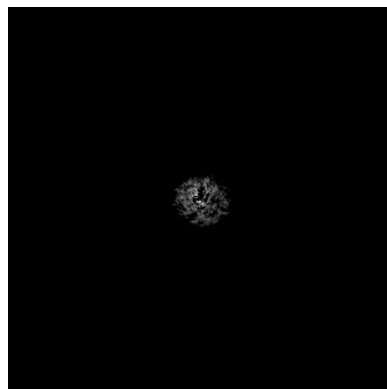
$$\text{error}(\mu) = \frac{\sum (|F_{uv}| c_{uv}^{1/2} - N_{uv}^{1/2})^2}{\sum_{u,v} N_{uv}}$$

(widely used error in optics)

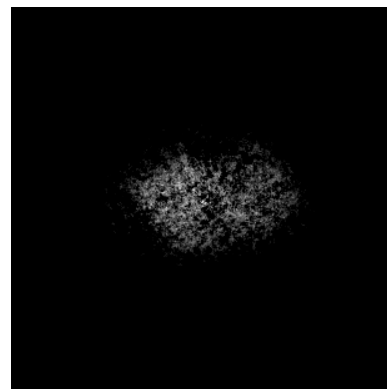
# SPR



(a) Reconstructed electron density with  $\mu = 10000$ .



(b) Reconstructed electron density with  $\mu = 100$ .



(c) Reconstructed electron density with  $\mu = 1$ .

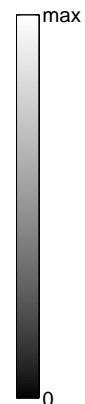


Figure: SPR results with different  $\mu$  for true diffraction data.

# SPR

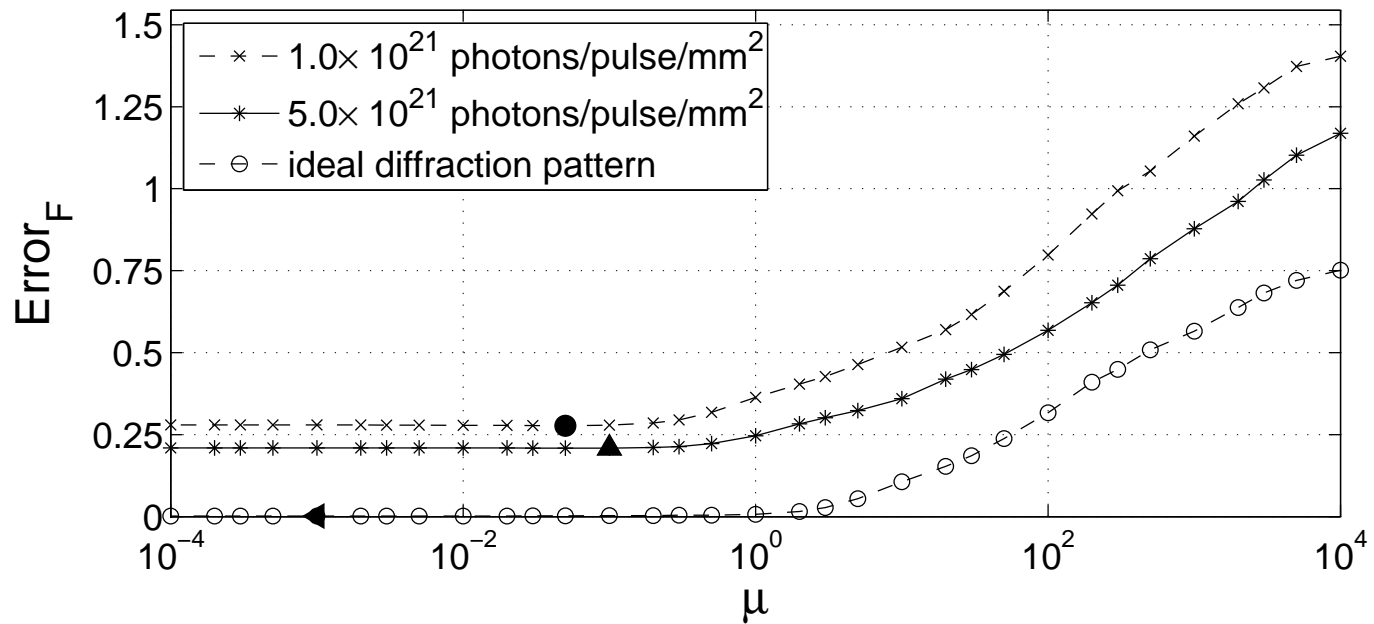
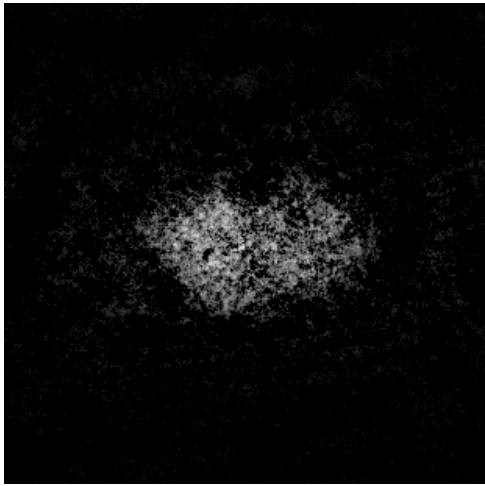
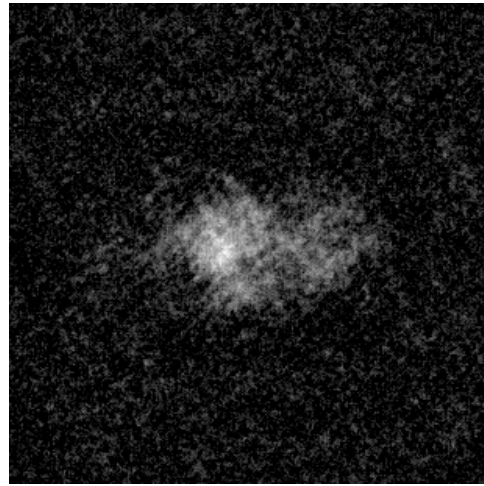


Figure: Error and  $\mu$ .

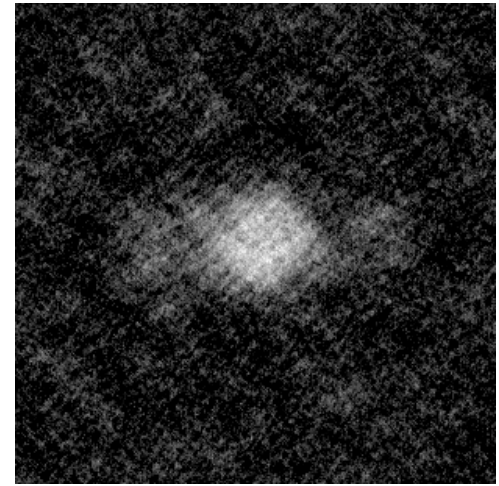
# SPR



(a) Reconstruction from diffraction image Fig.2b by the SPR method with  $\mu = 0.001$ .  $\text{Error}_F$  is  $1.60 \times 10^{-3}$ .



(b) Reconstruction from diffraction image Fig.2c by the SPR method with  $\mu = 0.1$ .  $\text{Error}_F$  is 0.209.



(c) Reconstruction from diffraction image Fig.2d by the SPR method with  $\mu = 0.05$ .  $\text{Error}_F$  is 0.277.

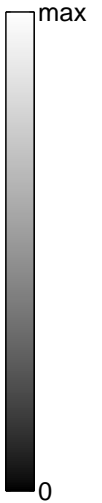
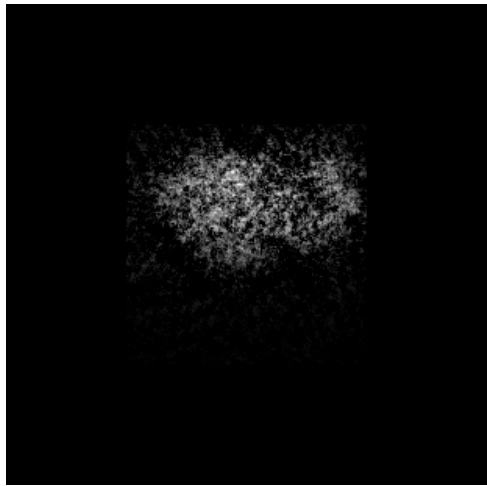
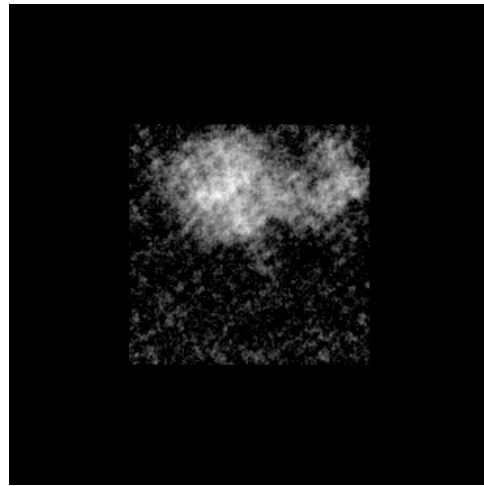


Figure: Reconstruction with SPR.

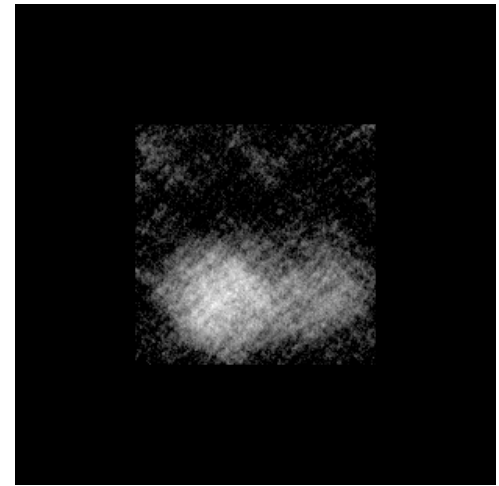
# HIO



(a) Reconstruction from diffraction image Fig.2b by the HIO method.  $Error_F$  is  $3.30 \times 10^{-3}$ .



(b) Reconstruction from diffraction image Fig.2c by the HIO method.  $Error_F$  is 0.220.



(c) Reconstruction from diffraction image Fig.2d by the HIO method.  $Error_F$  is 0.291.

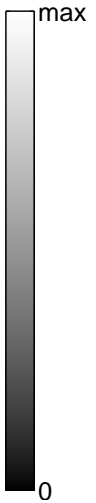
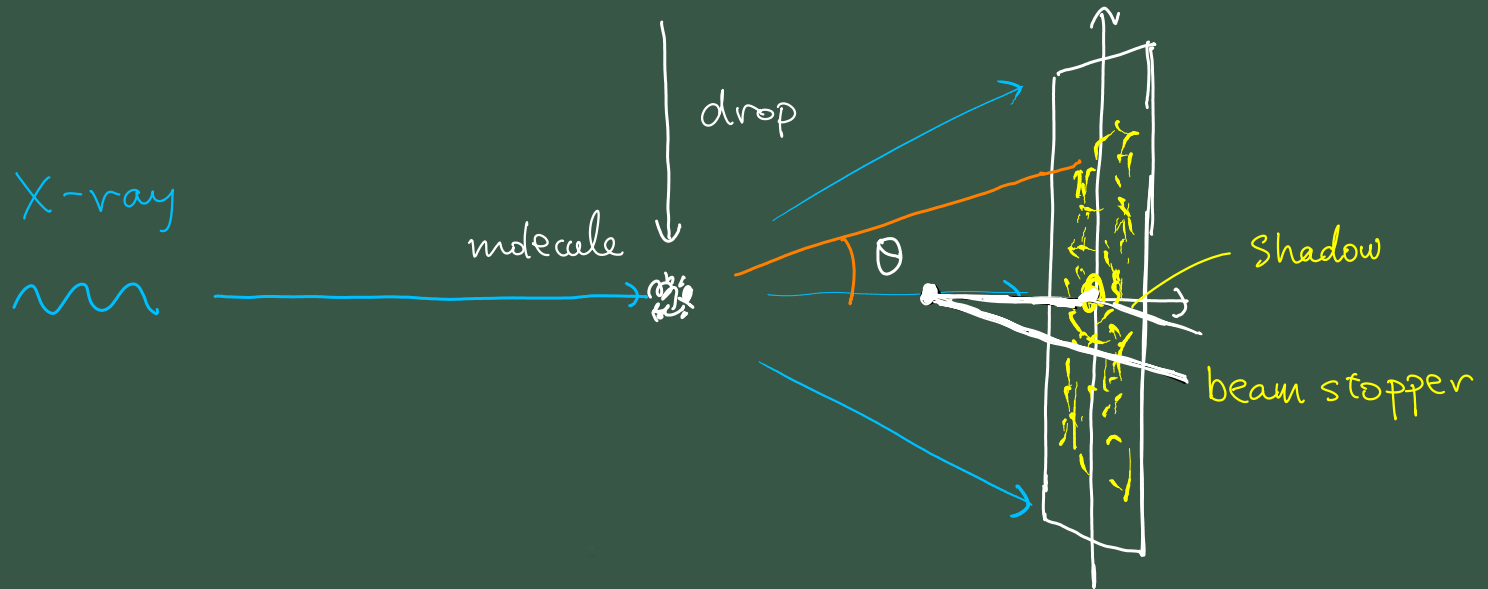


Figure: Reconstruction with HIO.

# Another Problem



There is a beam-stopper around the center.

308 x 308 diffraction image of Lysozyme

1 center pixel	→	5%	) photons are blocked
3x3	→	35%	
5x5	→	54%	

$$Q(f|N) = \underbrace{\sum_{uv \in A} (N_{uv} \ln |F_{uv}|^2 - |F_{uv}|^2 c_{uv})}_{\text{Likelihood}} - \underbrace{\mu \sum_{xy} W_{xy} f_{xy}}_{\text{prior}}$$

$A$ : visible pixel set.

- The same algorithm can be used.
- More robust against photons loss than the HIO method.



# SPR without 1 center pixel

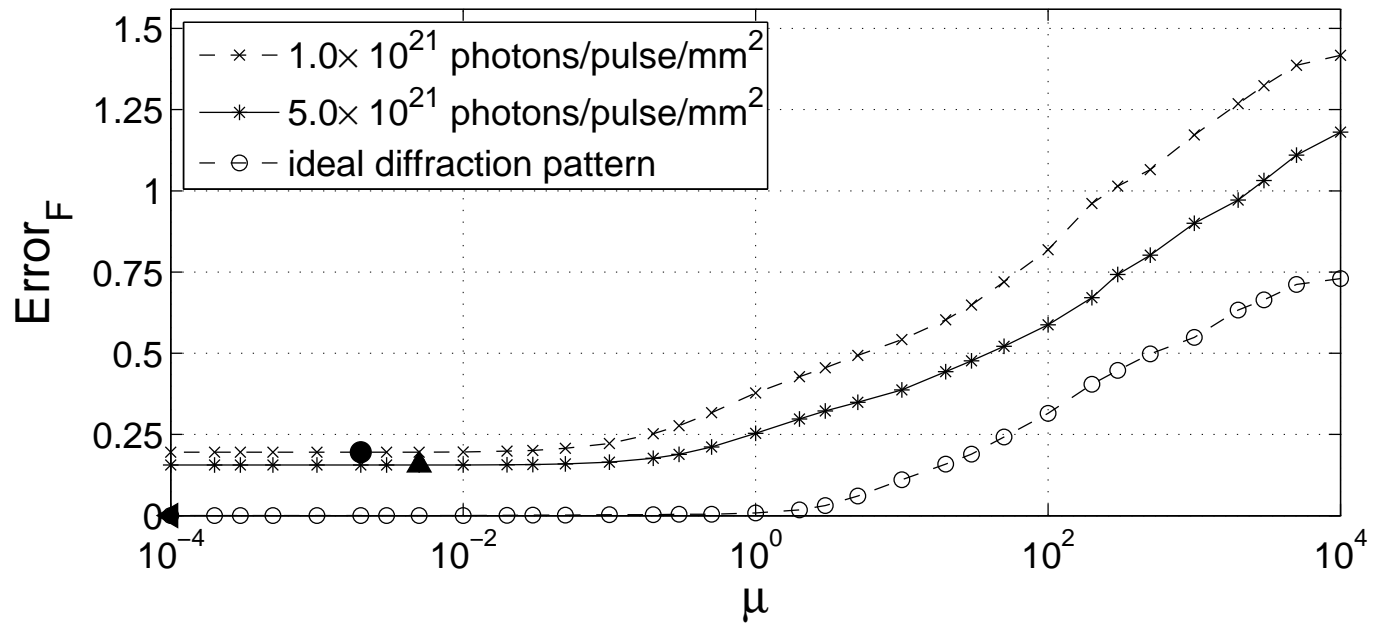
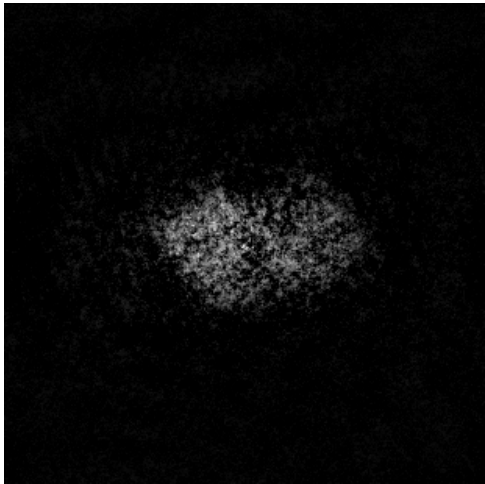
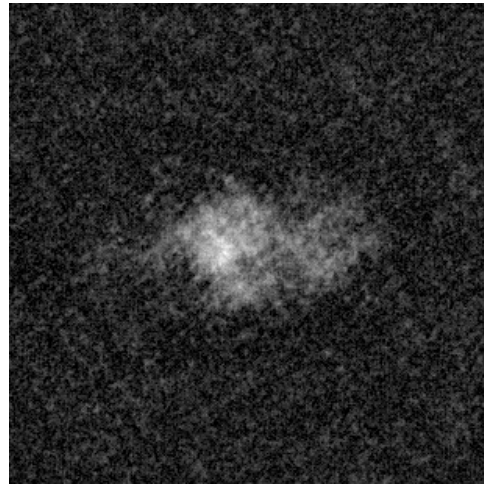


Figure: Error and  $\mu$ .

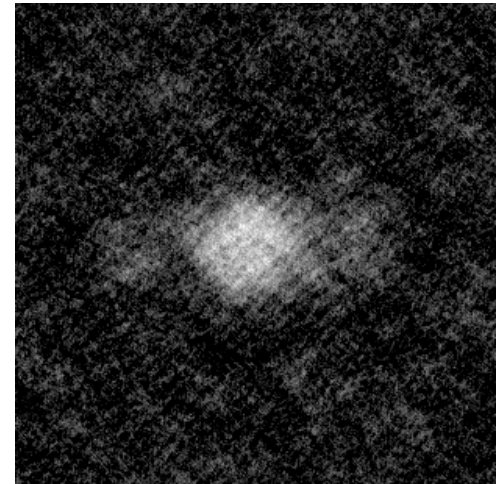
## SPR without 1 center pixel



(a) Reconstruction from diffraction image Fig.2b without the central pixel by the SPR method with  $\mu = 0.0001$ .  $Error_F$  is  $1.78 \times 10^{-6}$ .



(b) Reconstruction from diffraction image Fig.2c without the central pixel by the SPR method with  $\mu = 0.002$ .  $Error_F$  is 0.156.



(c) Reconstruction from diffraction image Fig.2d without the central pixel by the SPR method with  $\mu = 0.005$ .  $Error_F$  is 0.195.

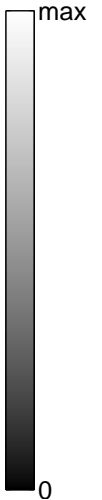
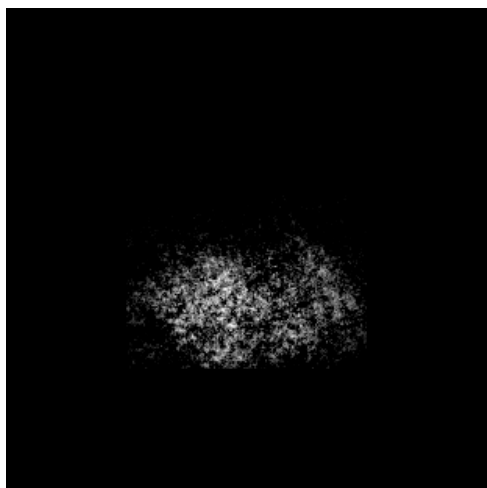
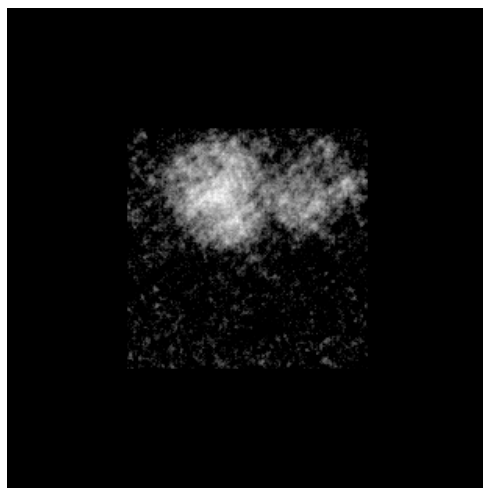


Figure: Reconstruction with SPR.

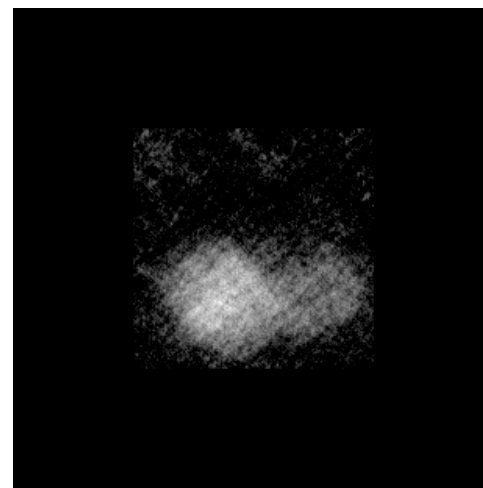
## HIO without 1 center pixel



(a) Reconstruction from diffraction image Fig.2b without the central pixel by the HIO method.  $Error_F$  is  $1.01 \times 10^{-2}$ .



(b) Reconstruction from diffraction image Fig.2c without the central pixel by the HIO method.  $Error_F$  is 0.245.



(c) Reconstruction from diffraction image Fig.2d without the central pixel by the HIO method.  $Error_F$  is 0.318.

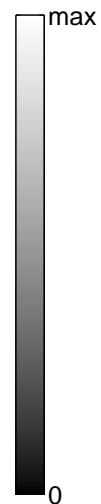


Figure: Reconstruction with HIO.

# SPR without $3 \times 3$ center pixels

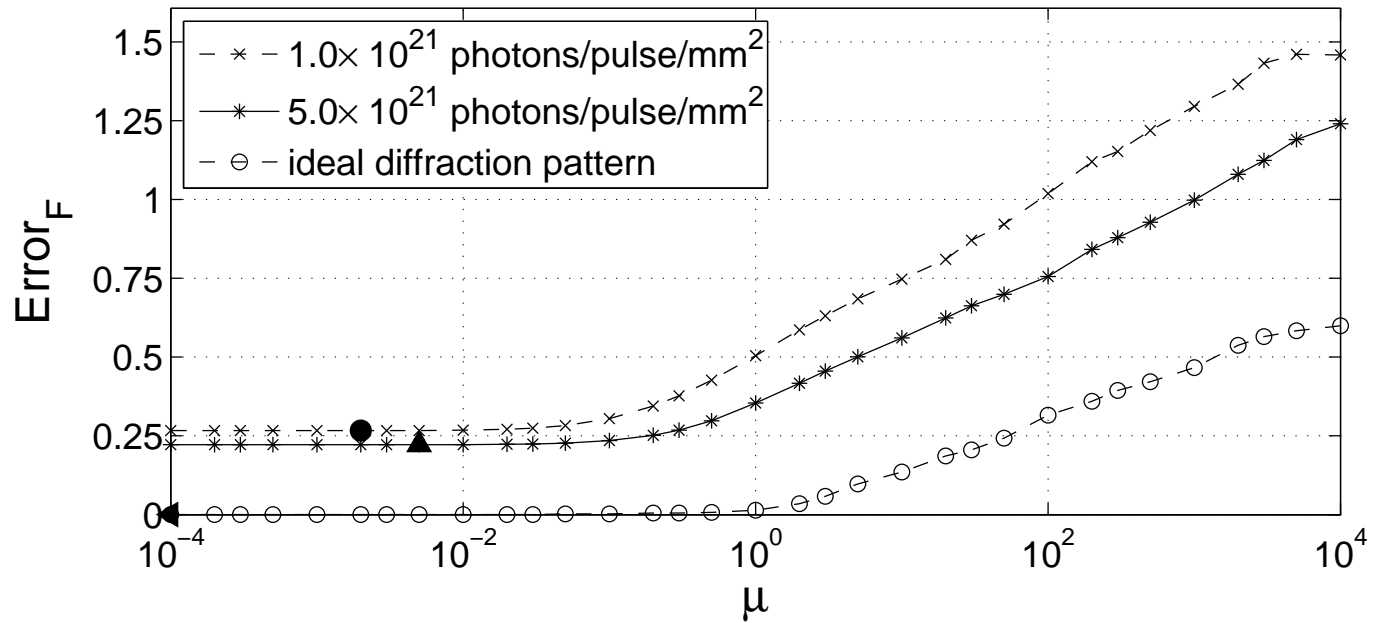
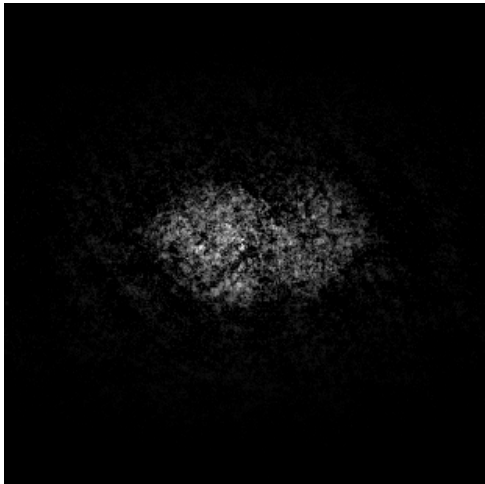
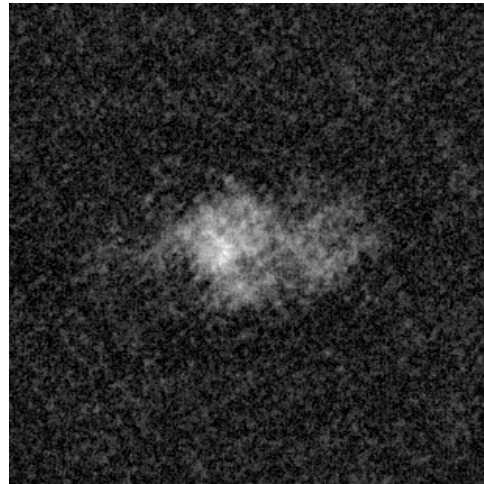


Figure: Error and  $\mu$ .

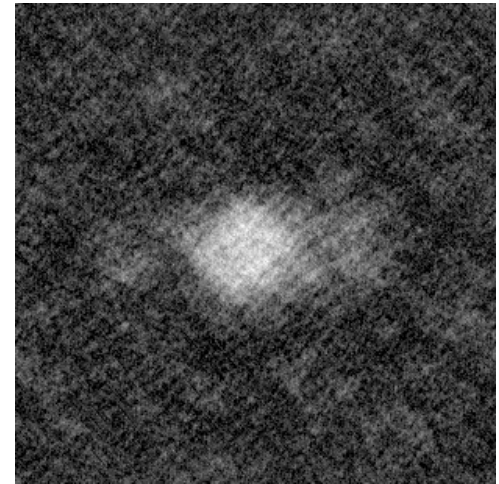
## SPR without $3 \times 3$ center pixels



(a) Reconstruction from diffraction image Fig.2b without the central  $3 \times 3$  pixels by the SPR method with  $\mu = 0.0001$ .  $\text{Error}_F$  is  $2.34 \times 10^{-6}$ .



(b) Reconstruction from diffraction image Fig.2c without the central  $3 \times 3$  pixels by the SPR method with  $\mu = 0.002$ .  $\text{Error}_F$  is 0.222.



(c) Reconstruction from diffraction image Fig.2d without the central  $3 \times 3$  pixels by the SPR method with  $\mu = 0.005$ .  $\text{Error}_F$  is 0.267.

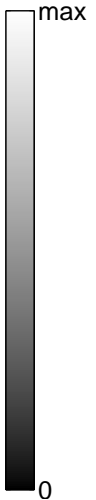
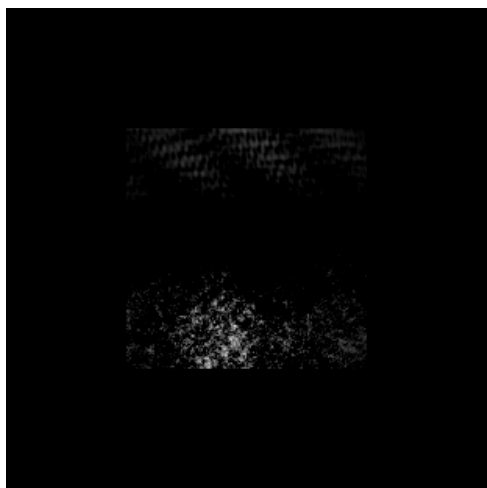
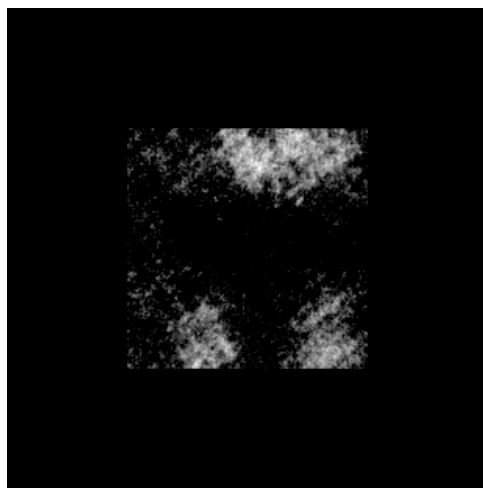


Figure: Reconstruction with SPR.

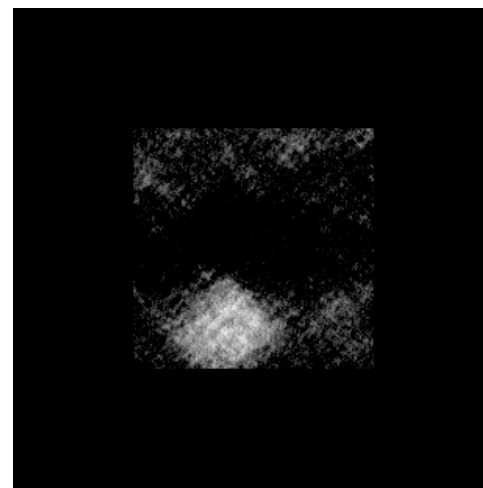
## SPR without $3 \times 3$ center pixels



(a) Reconstruction from diffraction image Fig.2b without the central  $3 \times 3$  pixels by the HIO method.  $Error_F$  is 0.128.



(b) Reconstruction from diffraction image Fig.2c without the central  $3 \times 3$  pixels by the HIO method.  $Error_F$  is 0.487.



(c) Reconstruction from diffraction image Fig.2d without the central  $3 \times 3$  pixels by the HIO method.  $Error_F$  is 0.557.

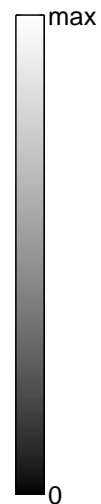


Figure: Reconstruction with HIO.

# Conclusion

- ① Proposed a new SPR method for phase retrieval
  - ① Based on sparsity of electron density.
  - ① Works well with small number of photons
  - ① Works well even some center pixels are blocked.
  - ① SPR method can be used instead of HIO.

# Compton Camera Imaging

joint work with

H. Odaka<sup>\*</sup>, M. Uemura<sup>†</sup>

T. Takahashi<sup>†</sup>, S. Watanabe<sup>†</sup>

& T. Takeda<sup>\*</sup>

† JAXA

† Hiroshima Univ.



① Compton camera

② Compton camera imaging

★ measurement process.

★ estimation method

★ improvement

③ Conclusion

## Q Compton Camera

Q Visualize  $\gamma$ -rays

Wide range of applications

- Astronomy (ASTRO-H)
- Medical application
- Visualize the contamination of soil (Fukushima)

# Astro-H project.

http://astro-h.isas.jaxa.jp/ ASTRO-H 次期X線国際天...

ファイル(E) 編集(E) 表示(V) お気に入り(A) ツール(I) ヘルプ(H) x 変換 選択

HOME ニュース・イベント ASTRO-Hの挑戦 スケジュール/開発日誌 画像ギャラリー 研究者向け情報 開発メンバーと関連サイト

New exploration X-ray Telescope English チームメンバー専用

# ASTRO-H

次期X線国際天文衛星

## 最新ニュース

ASTRO-Hが埼玉新聞で紹介されました。  
…詳細を読む  
2013年04月17日 [一般向け]

ASTRO-H日誌が公開されました。  
…詳細を読む  
2011年12月01日 [一般向け]

日本物理学会・日本天文学会で多くの発表が行われました。  
…詳細を読む  
2011年9月28日 [研究者向け]

X線天文学とASTRO-Hの特集がJAXAの機関誌「JAXA's」に掲載されました。  
…詳細を読む  
2011年5月1日 [一般向け]

ASTRO-Hの特集がJAXA HPIに掲載されました。  
…詳細を読む  
2011年4月14日 [一般向け]

## ASTRO-Hの挑戦

ASTRO-Hが拓く新しい宇宙像 ▶  
宇宙の謎にもう少し手が届く！

### プロジェクトについて

- ▶ 日本のX線天文衛星の歴史
- ▶ ASTRO-Hの概要

### X線天文学の世界 ▶

### ASTRO-Hのしくみ

- ▶ X線を集める—X線望遠鏡
- ▶ 画像を撮る—X線CODカメラと硬X線撮像検出器
- ▶ エネルギーを測る—マイクロカロリメーター

### 【イベント情報】

ASTRO-Hの挑戦を知って頂くためのイベントなどを、紹介しています

## 研究者向け情報

- ▶ 研究者向けニュース
- ▶ 研究会・シンポジウム
- ▶ 論文リスト
- ▶ シミュレーション用ツール
- ▶ 技術資料

## スケジュール/開発日誌

ASTRO-H第2回コラボレーション会議が開催されました。  
2010年2月23-25日

## 画像ギャラリー

# smartphone with a radiation detector

## SoftBank Pantone 5 107SH hands-on: radiation detection comes to Android

By Sam Byford on May 29, 2012 02:24 am [Email](#) [@345triangle](#)

DON'T MISS STORIES *FOLLOW THE VERGE* [g+](#) [Like](#) 243k [Follow](#) 335K followers



16 UPDATES TO

### Japan's summer 2012 phone lineup

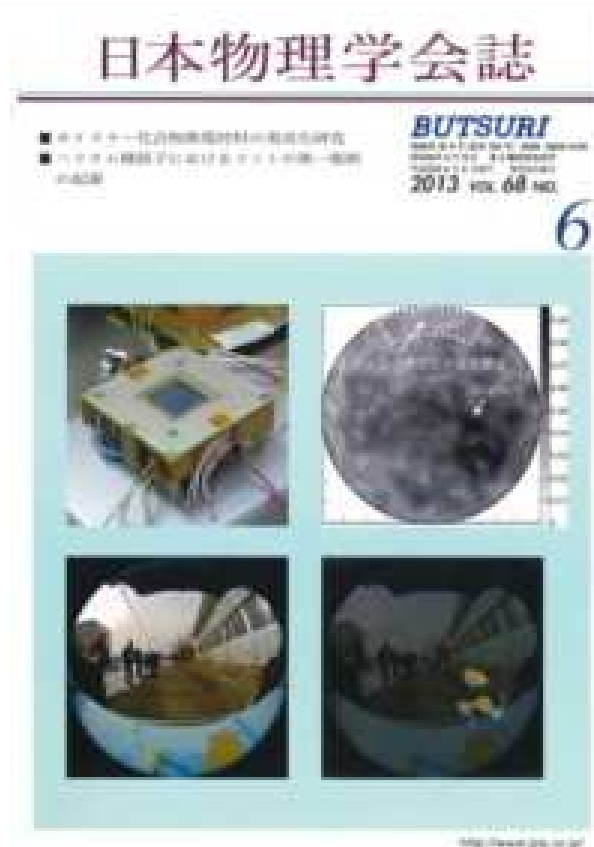
**JUN 19** Frog redesigns the Android user experience with Feel\_UX

**MAY 29** SoftBank's summer 2012 Android lineup: better signal, faster downloads

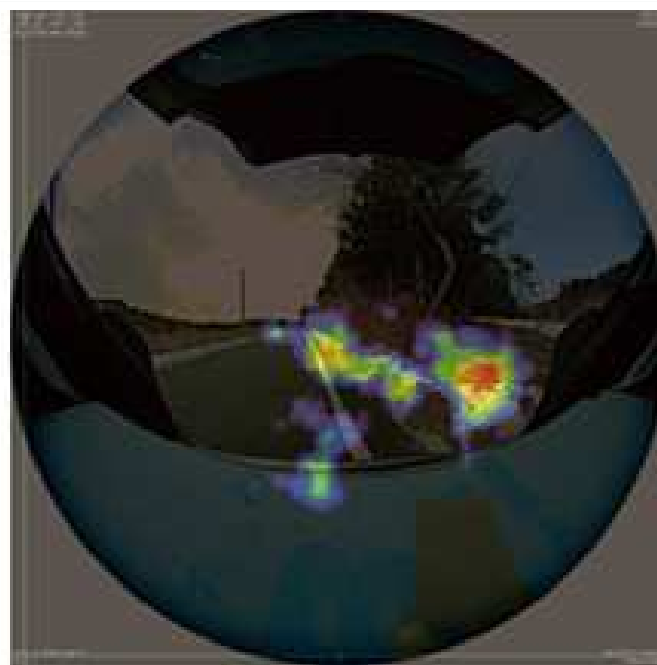
**MAY 29** 'Fastest ever' 110Mbps SoftBank 4G mobile router out in Japan this fall

SoftBank's Pantone 5 107SH will make headlines in the global press for one reason and

福島での応用 (JAXA, 日本物理学会の HP より).



福島での応用(JAXA の HP より).



①  $\gamma$ -rays: Electromagnetic radiation with wavelength ( $< 10\text{pm}$ )

- Difficult to detect.

- Measurement techniques

- Geiger-counter

- Scintillation detector

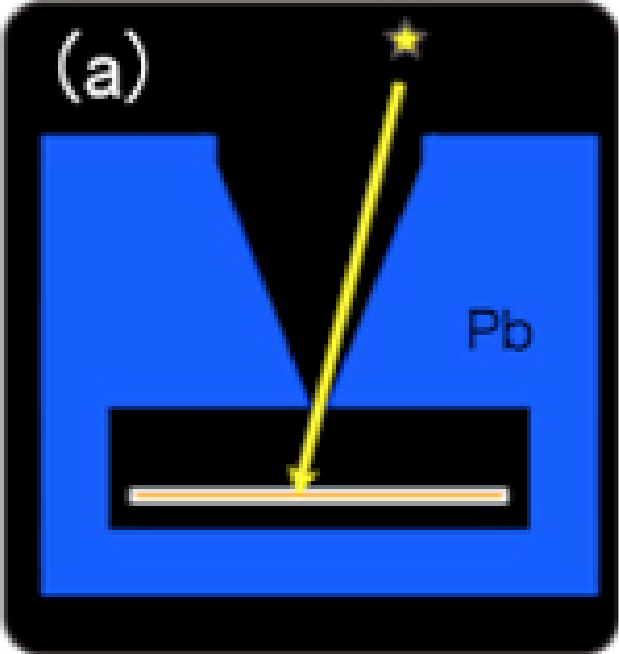
) cannot measure direction of arrival

- Compton camera

---- use Compton scattering

can be used for imaging

gamma-ray camera as a pinhole camera





Q Compton camera

Q Compton camera imaging

☆ measurement process.

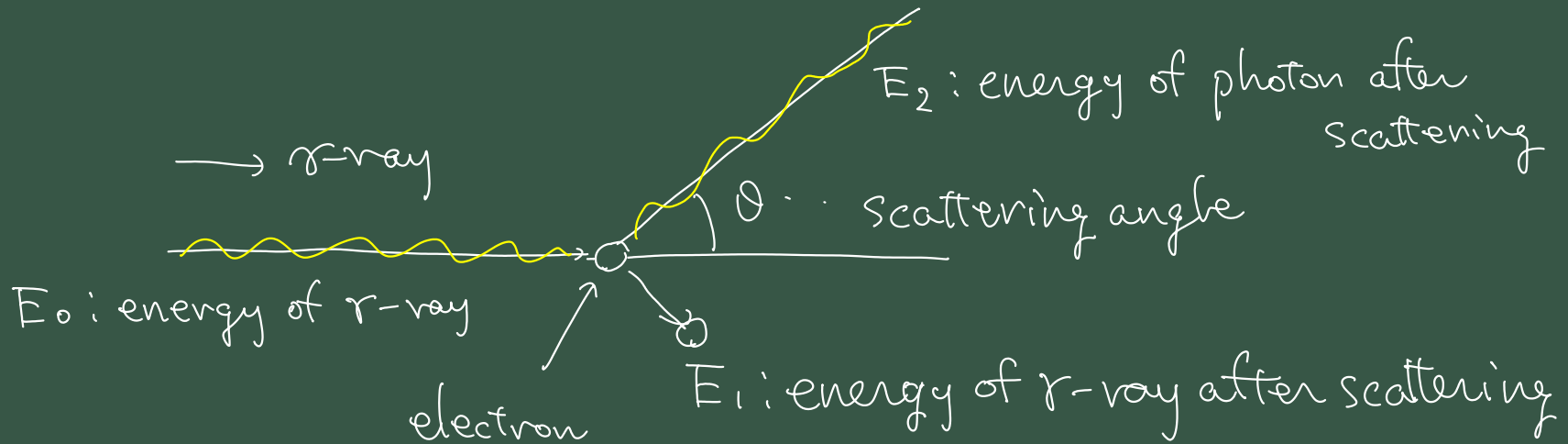
☆ estimation method

☆ improvement

Q Conclusion

# Compton camera imaging

## Compton scattering

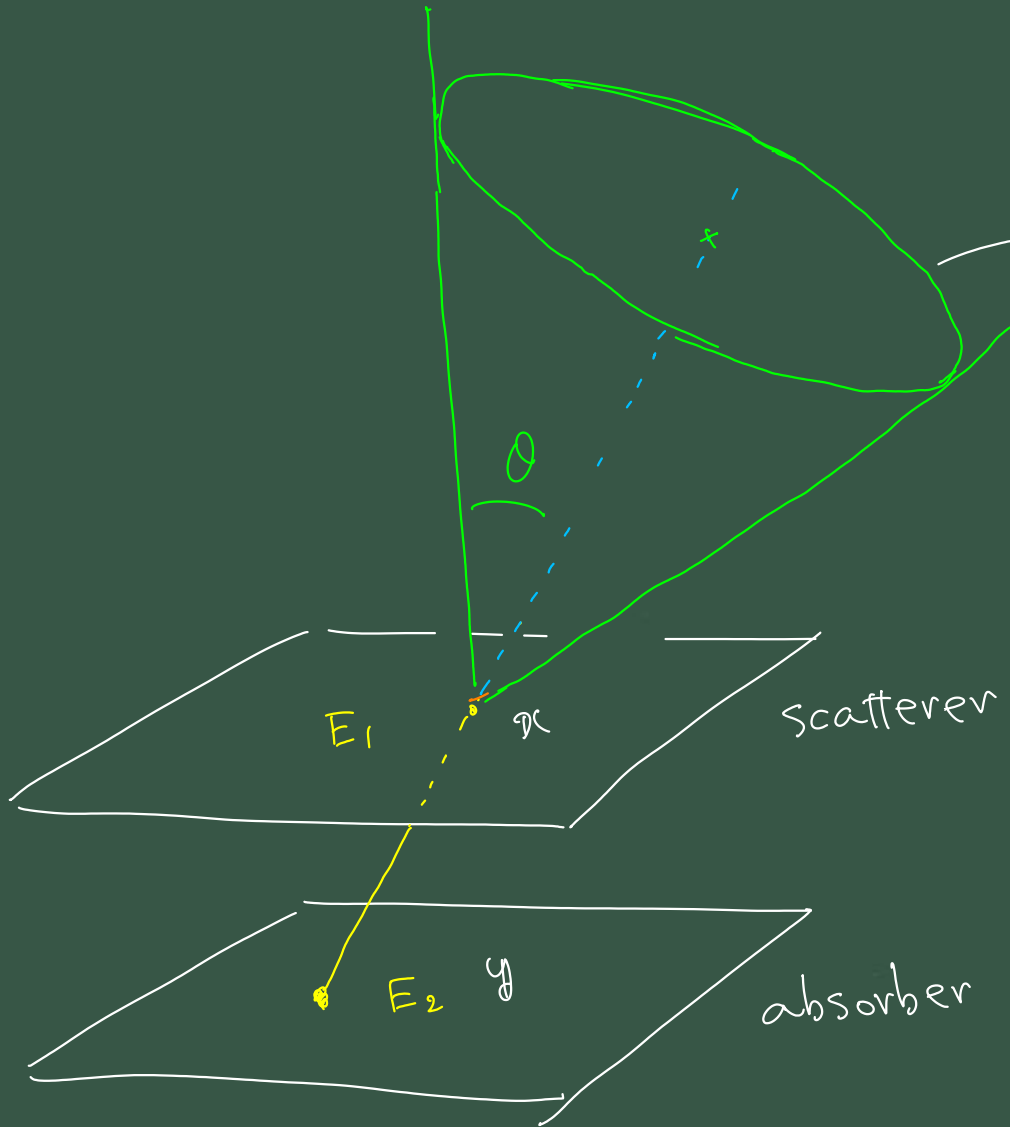


$$\underline{\frac{1}{E_2} - \frac{1}{E_0}} = \underline{\frac{1}{E_2} - \frac{1}{E_1 + E_2}} = \frac{1}{m_e c} (1 - \cos \theta)$$

$m_e$  : mass of electron

$c$  : speed of light

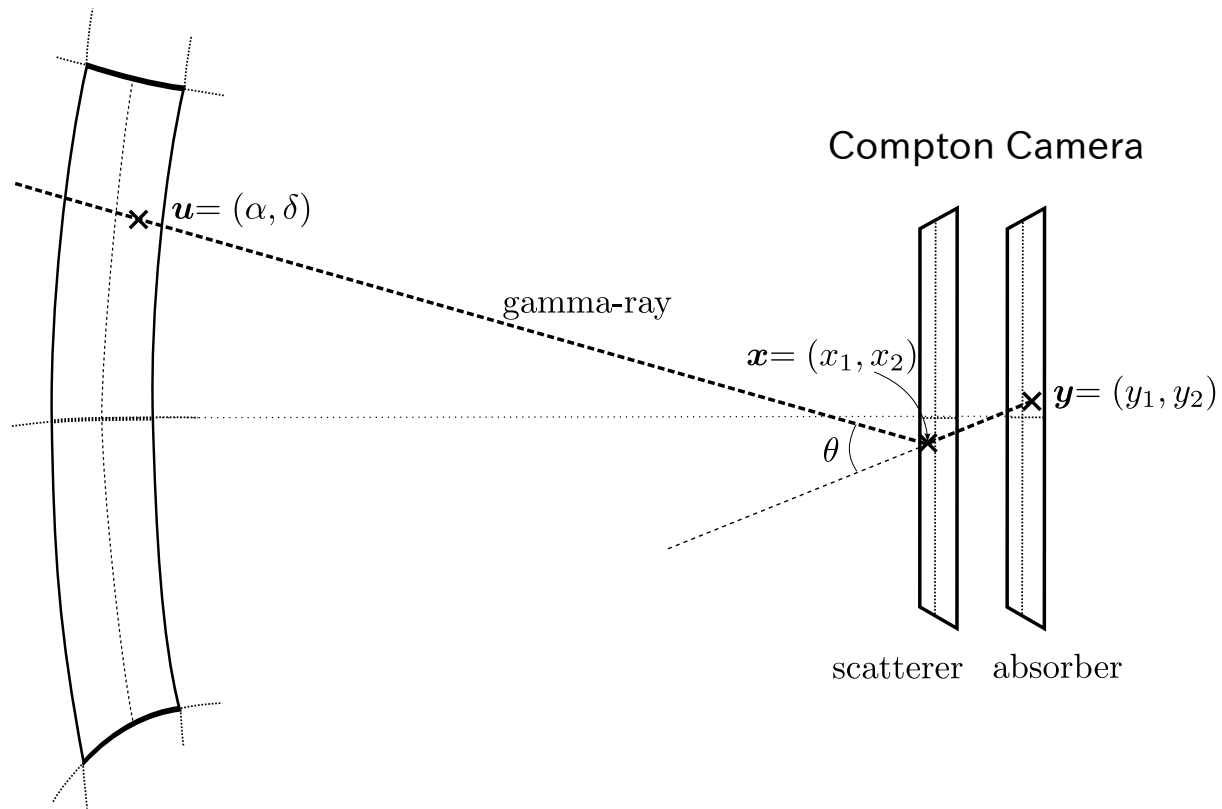
# Compton camera imaging



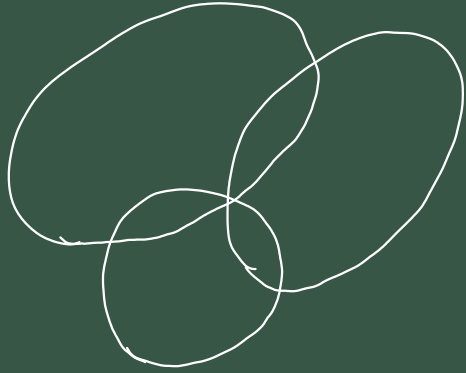
Photon came from a point on this cone

Observing many photons and  $\gamma$ -ray intensity map can be reconstructed.

# Compton camera imaging.



② Conventional method for Compton camera imaging.



② Back projection.

- Draw an ellipsoid for each observation
- Simple and fast



Image becomes blurred.

## Physical parameters of the simulated semi-conductor.

		density [g cm <sup>-3</sup> ]	energy resolution FWHM[keV]
scatterer	Si	2.33	2.0
absorber	CdTe	5.86	2.0

Q Compton camera

Q Compton camera imaging

☆ measurement process.

☆ estimation method

☆ improvement

Q Conclusion

# ① Probabilistic framework for Compton camera imaging

$\gamma$ -ray

- DOA :  $u$
- $\gamma$ -ray strength from  $u$  :  $\lambda(u)$

Compton camera

- Probability of scattering :  $s(u)$   
position of scattering on scatterer :  $x$   
= absorption on absorber :  $y$

- For large distance sources:  $w = x - y$

- Scattering angle :  $\cos \theta$

- model :  $p(w | u)$

$$\rightarrow \underline{w = (z, \cos \theta)}$$

(Scattering angle  $\theta$  follows  
Klein-Nishina formula)



## ☆ Observation Process

$Y(u)$ : The number of photons observed at position  $u = (\omega, \cos\theta)$  follows a Poisson dist.

$$\underline{Y(u) \sim \text{Poisson} \left( \sum_{u'} p(u|u') \lambda(u') s(u') \right)}$$

re-parameterization

$$q(u) = \frac{Y(u)}{\sum_{u'} Y(u')}, \quad p(u) = \frac{\lambda(u) s(u)}{\sum_{u'} \lambda(u') s(u')}$$

$$\underline{q(u) = \sum_{u'} p(u|u') p(u')}$$

$q(u)$ : Prob. that absorbed photons is observed at  $u$

$p(u)$ :  $\stackrel{=}{}$  came from  $u$

Q  $p(\nu | \omega)$ ,  $S(\omega)$

Q Theoretically, Klein-Nishina formula can be used to compute  $p(\nu | \omega)$ , but we use a Monte Carlo simulation

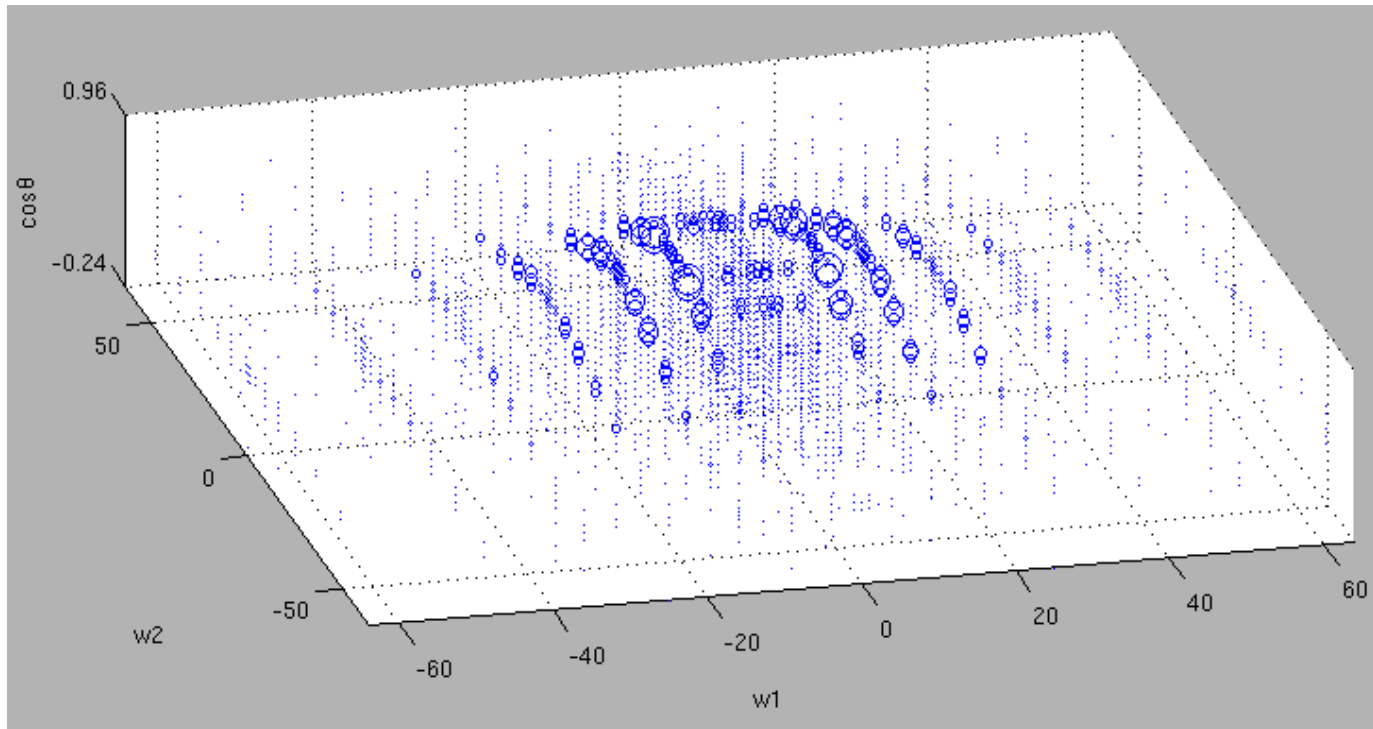
Q Well-designed physical simulation with real camera can be used to measure  $p(\nu | \omega)$  physically.

Q  $p(\nu | \omega)$  and  $S(\omega)$  must be prepared for each energy level  $E_0$ . This should be done off-line beforehand.

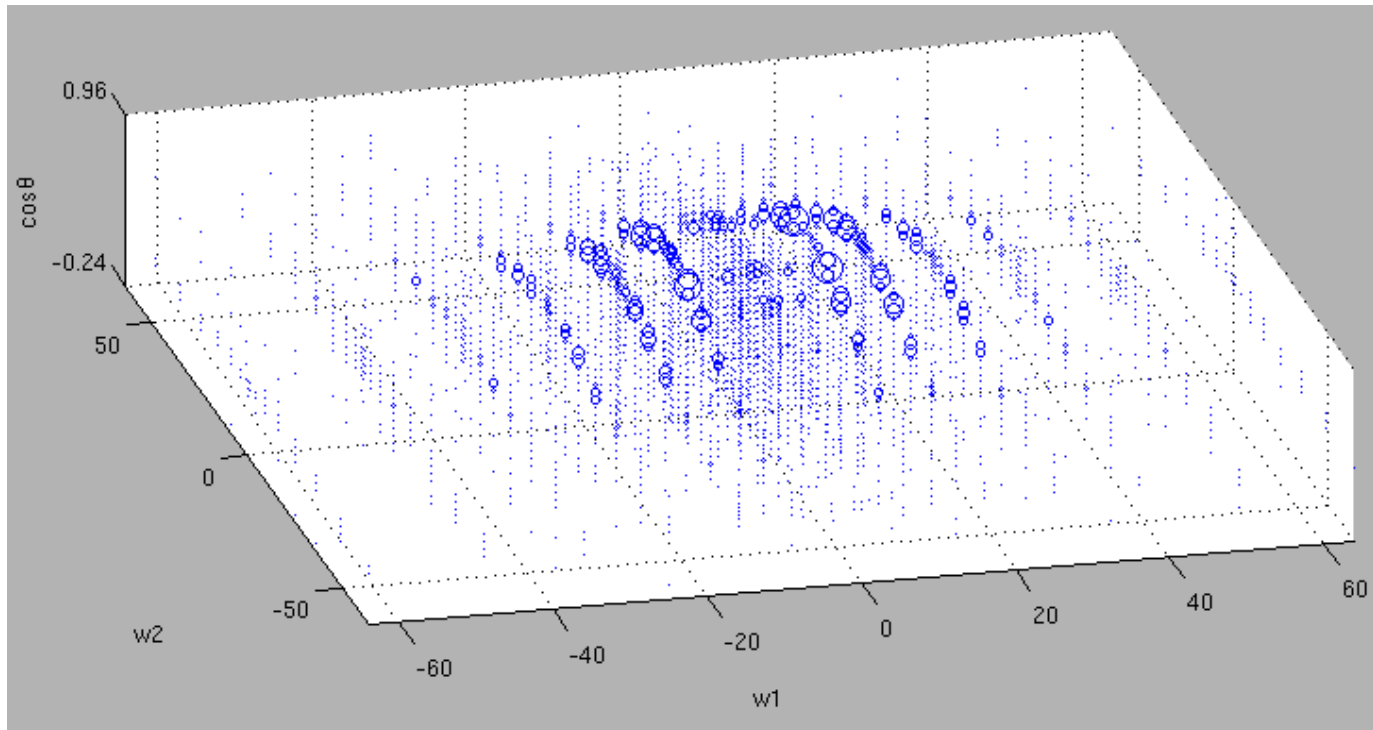
Design of bins of  $u$  and  $v$ .

	$u$		$v$		$\cos \theta$
	$\alpha$	$\delta$	$\frac{w_1}{\sqrt{ w_1 }}$	$\frac{w_2}{\sqrt{ w_2 }}$	
	[degree]		[mm <sup>1/2</sup> ]		
min	-30	-30	-8.2	-8.2	-.24
max	30	30	8.2	8.2	.96
# of bins	21	21	17	17	24

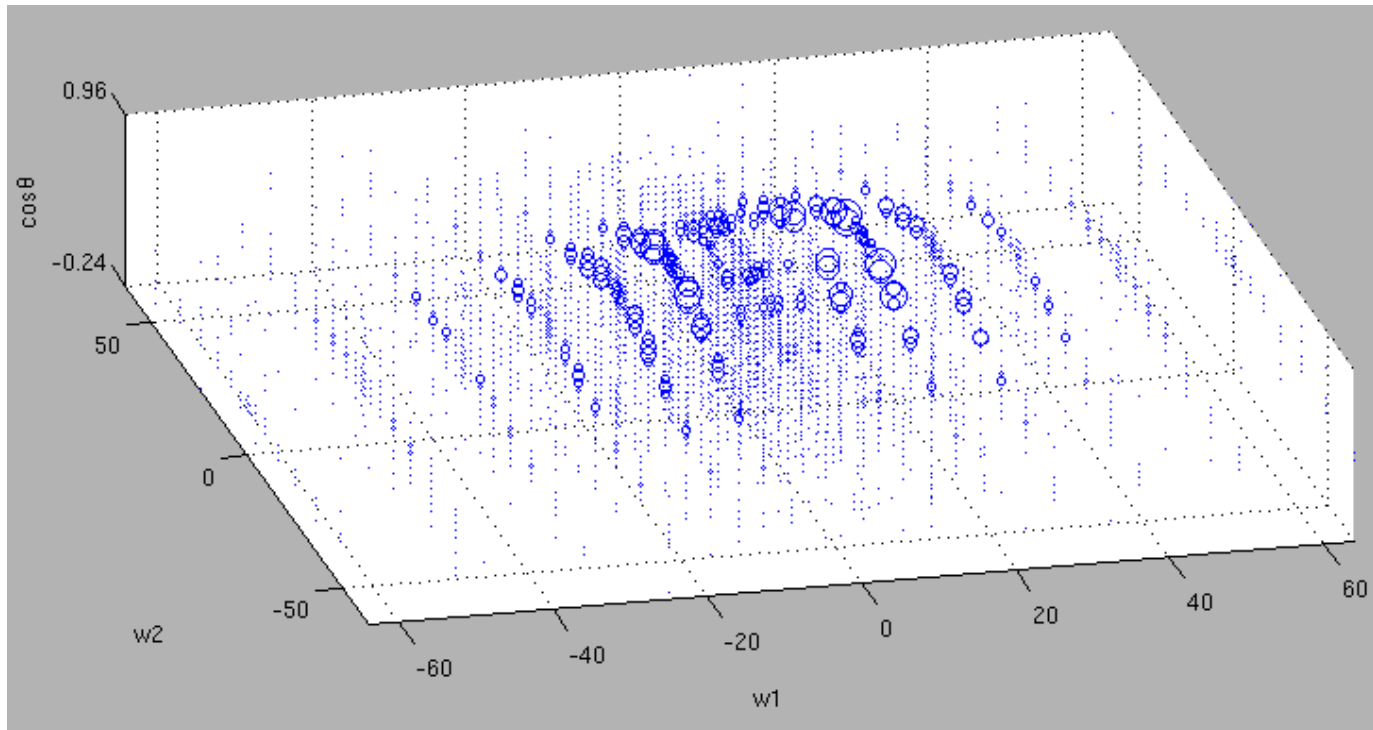
Distribution of the received photon from center.



Distribution of the received photon from  $8^\circ$  off center.



Distribution of the received photon from  $12^\circ$  off center.



## ☆ Estimation method

$$\sum_u q(u) = 1, \quad \sum_u p(u) = 1, \quad \sum_v p(v|u) = 1,$$

mixture dist.  $q(v)$  =  $\sum_u$   $\overset{\text{given}}{p(v|u)}$   $\overset{\text{dist. to estimate}}{p(u)}$

Samples from a mixture dist. are observed.

Each mixture component is known ( $p(v|u)$ )

we want to estimate the mixing coefficient ( $p(u)$ ).

## Q Estimation method (EM algorithm)

Observation:  $v_x$  ( $x=1, \dots, N$ )

Starting from  $p^{(0)}(\omega)$ , update  $p^{(t)}(\omega)$  as follows.

E-step

$$q^{(l)}(v_x) = \sum_{\omega} p(v_x | \omega) p^{(l)}(\omega)$$

M-step

$$p^{(l+1)}(\omega) = \frac{1}{N} \sum_{x=1}^N \frac{p(v_x | \omega)}{q^{(l)}(v_x)} p^{(l)}(\omega)$$

---

Two problems

Q  $\hat{p}(\omega)$  is not sparse  $\rightarrow$  MAP estimate with Dirichlet prior

Q Slow convergence  $\rightarrow$  Approximate Fisher's scoring.



Q Compton camera

Q Compton camera imaging

☆ measurement process.

☆ estimation method

☆ improvement

Q Conclusion

## Q MAP estimation (Bayesian approach)

In astronomy approximation,  $p(u)$  should be sparse (lot of 0's).

$p(u)$  is a multinomial dist. Its conjugate prior is Dirichlet dist.

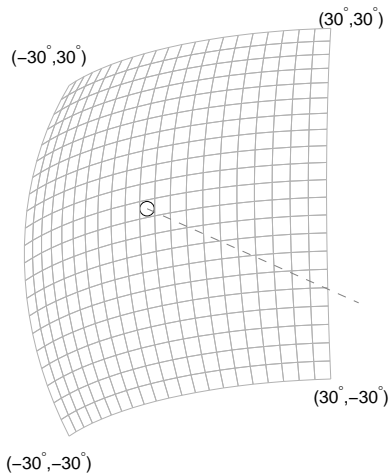
$$\pi_{\alpha}(p) = \frac{\Gamma(\alpha M)}{\Gamma(\alpha)^M} \prod_u p(u)^{\alpha-1}, \quad (M \text{ is the \# of } u)$$

MAP estimate

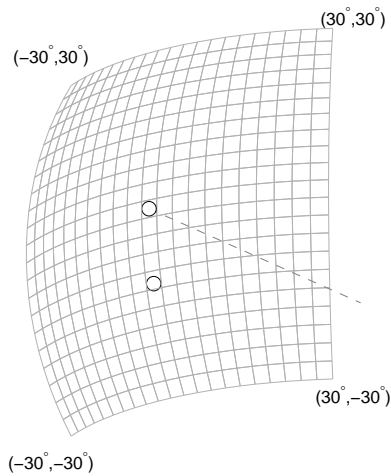
$$\begin{aligned} \hat{p}_{\text{MAP}} &= \operatorname{argmax} [\log \mathcal{P}(p | u_1, \dots, u_n)] \\ &= \operatorname{argmax} \left[ \underbrace{(\alpha-1) \sum_u \log p(u)}_{\text{prior distribution}} + \underbrace{L(p)}_{\text{log-likelihood}} \right] \end{aligned}$$

for  $\alpha < 1$ ,  $\hat{p}_{\text{MAP}}$  generally becomes sparse.

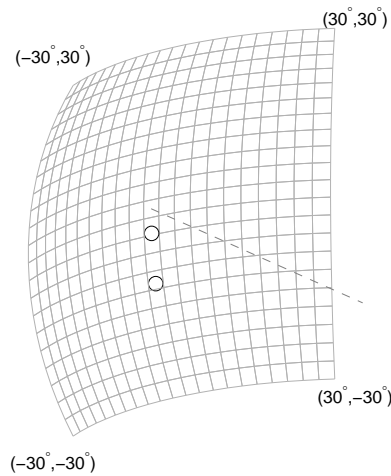
# Gamma-ray sources used for numerical simulations.



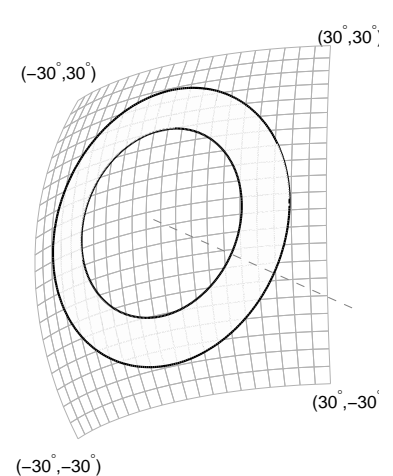
(a) single point source.



(b) two point source.

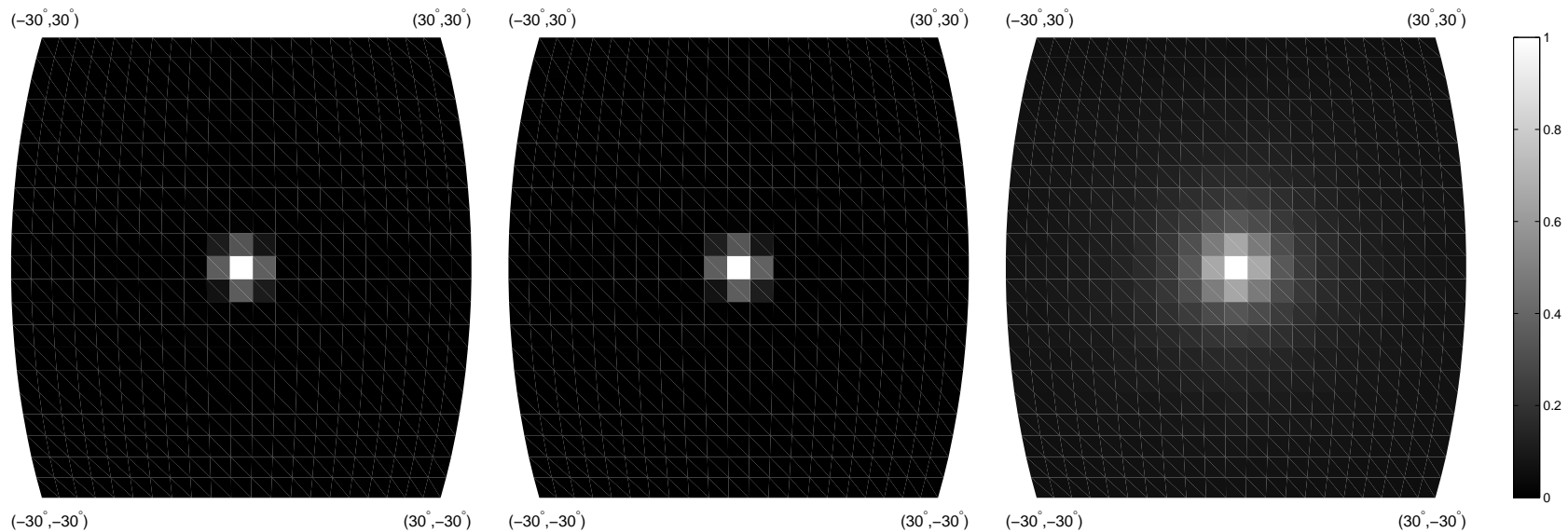


(c) two point source.



(d) distributed source.

# Reconstructed images for data from (a) single point source.

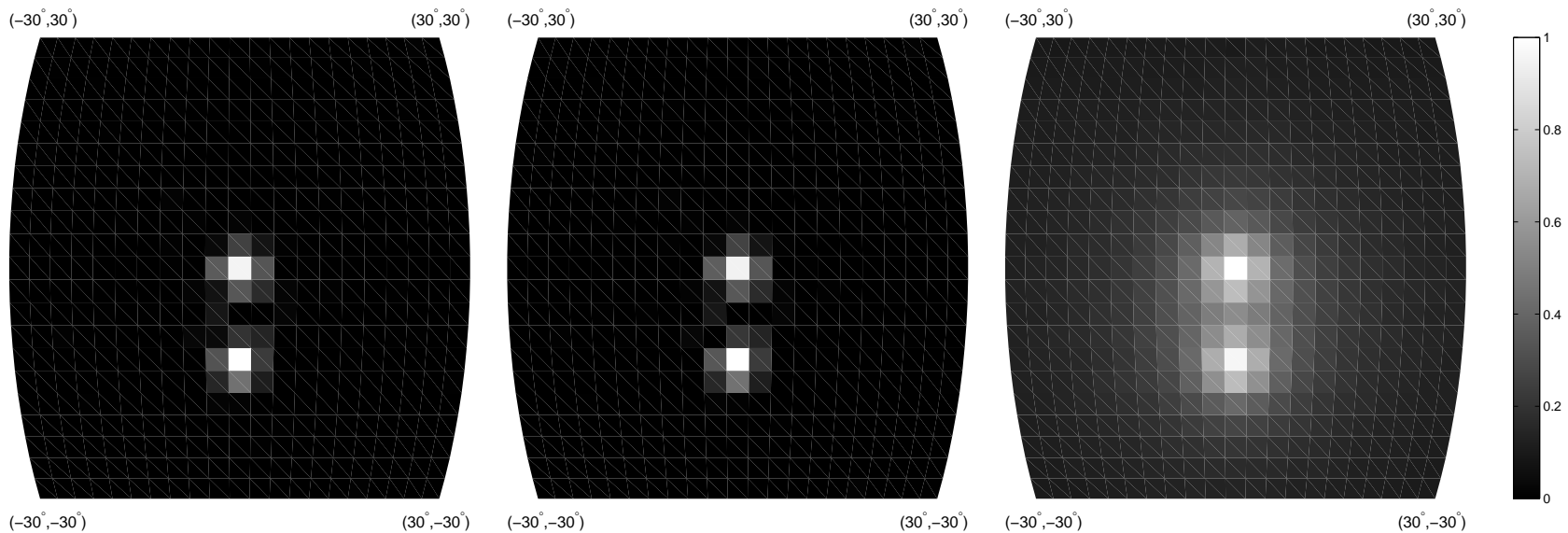


MLE reconstruction.

MAP reconstruction.

Back projection reconstruction.

# Reconstructed images for data from (b) two point source.

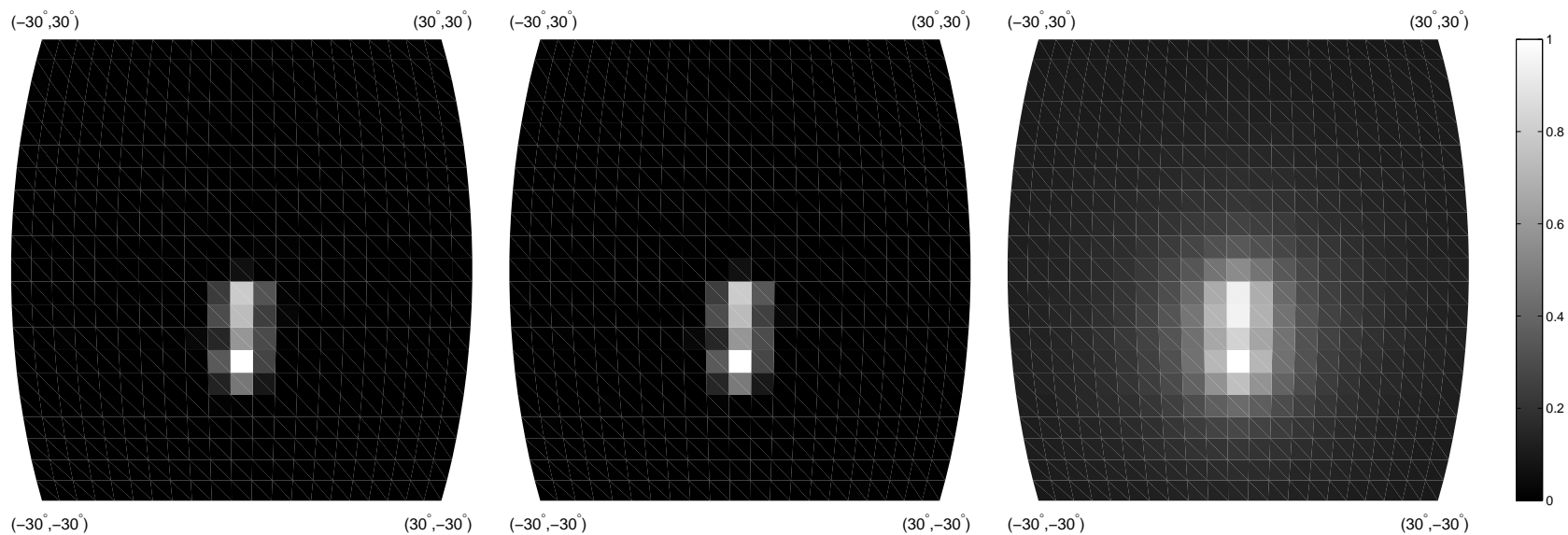


MLE reconstruction.

MAP reconstruction.

Back projection reconstruction.

# Reconstructed images for data from (c) two point source.

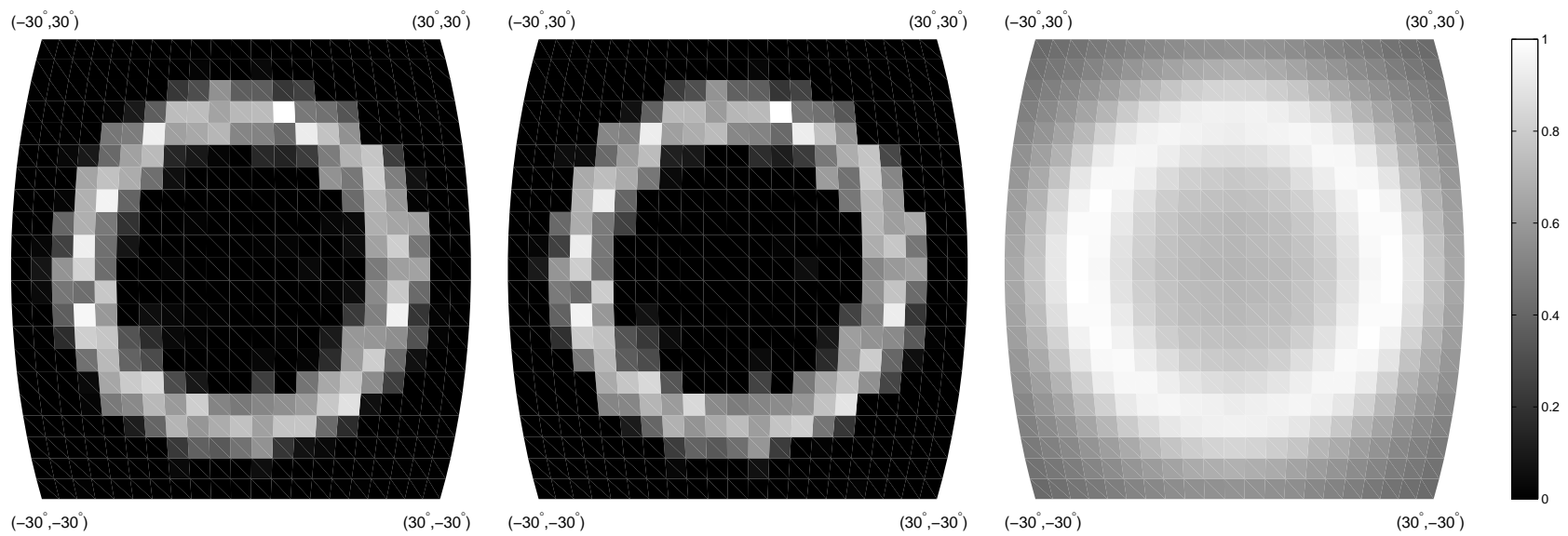


MLE reconstruction.

MAP reconstruction.

Back projection reconstruction.

# Reconstructed images for data from (d) distributed source.



MLE reconstruction.

MAP reconstruction.

Back projection reconstruction.

Q Compton camera

Q Compton camera imaging

☆ measurement process.

☆ estimation method

☆ improvement

Q Conclusion



## ① Conclusion

- ① Probabilistic framework for Compton camera imaging
- ① MAP estimation for sparse solution
- ① Speed up version of the EM algorithm.

## Ongoing projects

- ① multi-layered camera
- ① other applications, such as Fukushima soil contamination