# *M*-estimation of wavelet variance

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**Abstract** The wavelet variance provides a scale-based decomposition of the process variance for a time series or a random field and has been used to analyze various multiscale processes. Examples of such processes include atmospheric pressure, deviations in time as kept by atomic clocks, soil properties in agricultural plots, snow fields in the polar regions and brightness temperature maps of South Pacific clouds. In practice, data collected in the form of a time series or a random field often suffer from contamination that is unrelated to the process of interest. This paper introduces a scale-based contamination model and describes robust estimation of the wavelet variance that can guard against such contamination. A new *M*-estimation procedure that works for both time series and random fields is proposed, and its large sample theory is deduced. As an example, the robust procedure is applied to cloud data obtained from a satellite.

**Keywords** Hermite expansion · Multiscale processes · Pockets of open cells · Robust estimation · Time series analysis

# 1 Introduction

Wavelets decompose a stochastic process (e.g., a time series or a random field) into sub-processes, each one of which is associated with a particular scale. A wavelet variance is the variance of a sub-process at a given scale and quantifies the amount of

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variation present at that particular scale. Wavelet variances at different scales enable us to perform an analysis of variance (see Percival and Walden 2000 and the references therein). This variance is of interest in numerous geophysical and related applications, where the overall observed process is an ensemble of sub-processes that operate at different scales. In addition, the wavelet variance serves as an exploratory tool to study power law processes (Stoev and Taqqu 2003), detect inhomogeneity (Whitcher et al. 2002), estimate spectral densities indirectly (Tsakiroglou and Walden 2002), and handle processes that are locally stationary with time- and space-varying spectra (Nason et al. 2000). Applications include the analysis of textures (Unser 1995), electroencephalographic sleep state patterns of infants (Chiann and Morettin 1998), the El Niño-southern oscillation (Torrence and Compo 1998), soil variations (Lark and Webster 2001), solar coronal activity (Rybák and Dorotovič 2002), the relationship between rainfall and runoff (Labat et al. 2001), ocean surface waves (Massel 2001), surface albedo and temperature in desert grassland (Pelgrum et al. 2000), heart rate variability (Pichot et al. 1999) and the stability of the time kept by atomic clocks (Greenhall et al. 1999).

In practice, however, data collected in the form of a time series or random field often suffer from contamination, which can occur in various ways. For example, the satellite cloud data we consider in Sect. 8 can be contaminated in at least three ways: at the satellite itself, during transmission of the signal through the atmosphere, and due to the presence of aberrant cloud types in a region where a single cloud type is dominant. A moderate amount of contamination often has a very adverse effect on conventional estimates of the wavelet variance. In classical statistics, it is common to handle contamination via removing questionable data points and then applying conventional methods; however, locating contamination in a time series or in a random field is often complicated because of dependence in the data. A better approach is to consider a contamination model, and then derive a robust procedure that protects against that model. In such situations, one hopes to measure robustness via influence functions and to form estimates that reduce the effect of bias in an optimum way. Several contamination models have been suggested in the time series literature. These include pure replacement models, additive outliers, level shift models and innovation outliers that hide themselves in the original time series. Chapter 8 of Maronna et al. (2006) provides an excellent account of the theory and methods of robust time series statistics based on these types of contamination models. However, robust nonparametric estimates of second-order statistics based on these contaminated models still present difficult problems. For example, Robinson (1986) observes that, under both pure replacement and additive outlier models, autocovariances of the unobserved uncontaminated process are not distinguishable from those of the observed contaminated process. Robinson's critique extends to wavelet variances, making the estimation and the inference of wavelet variances for contaminated data a difficult problem. Moreover contamination can be quite complex when a time series or random field is multiscale in nature. Different contamination processes can act on different scales independently, but contamination from one scale can leak into another and be hard to detect. A practical approach is to use the median of the squared wavelet coefficients

rather than their mean to estimate the wavelet variance. Stoev et al. (2006) touch on the usefulness of median-type estimators to guard against contamination. In this paper, we develop a full *M*-estimation theory for the wavelet variance and derive its large sample theory when the underlying process is Gaussian. Our approach treats the wavelet variance as a scale parameter and offers protection mostly against scale-based multiplicative contamination (i.e., additive on the log of squared wavelet coefficients) that acts independently at different scales. We will return to the discussion on robustness in Sect. 9.

To illustrate our proposed methodology, let us consider the following example motivated by vertical shear measurements in the ocean (see Sect. 7 for details). Such series are subject to bursts of turbulence isolated in a select number of scales, but burstfree segments are well modeled by a stationary process. Figure 1a shows a simulated Gaussian series of length 4,096 from this stationary process. We compute its wavelet coefficients for scales  $\tau_i = 2^{j-1}$ , j = 1, ..., 9, and, following Percival (1995), form 95% confidence intervals (CIs) for the wavelet variance based upon averaging squared coefficients for a given scale (left-most lines in Fig. 2 in the set of 4 lines jittered about each of the 9 scales). We form an alternative estimator by taking the median of the squared coefficients and correcting for bias as dictated by our theory (see Eq. (15)). The second-from-left lines in each group of four show the 95% CIs based upon the robust median-based estimates. We see good agreement with the mean-type estimates at all scales. Next we contaminate the time series by introducing scale-based multiplicative noise, which is intended to simulate bursts of turbulence focused around scale  $\tau_3$ (Fig. 1b). The second-from-right lines in Fig. 2 show the CIs for the wavelet variance based upon this contaminated data and the usual mean-type estimator. Outliers significantly change the wavelet variance estimates at small scales. Any inferences that we might want to draw about the turbulent-free vertical shear process at small scales would be materially influenced by the contamination. Finally the right-most lines in



**Fig. 1** Simulated ocean shear data. The *top plot* **a** is of the original simulated series, while the *bottom plot* **b** shows the series after contamination



**Fig. 2** Confidence intervals (*CIs*) for the wavelet variance based upon mean- and median-type estimators of the wavelet variance using simulated ocean shear data (Fig. 1a) and a contaminated version thereof (Fig. 1b). CIs for nine scales  $\tau_j = 2^{j-1}$ , j = 1, ..., 9, are displayed. For each scale there are four lines jittered about the level *j*. From *left* to *right*, these show 95% CIs based upon the mean-type estimator with uncontaminated data; median-type with uncontaminated; mean-type with contaminated; and median-type with contaminated

Fig. 2 show CIs based upon our robust median-based estimator and the contaminated series. The CIs are close to those for the uncontaminated series, making it possible to draw inferences about the turbulent-free vertical shear process from the contaminated data. Our robust procedure thus works well when observed data are subjected to scale-based multiplicative contamination of an underlying Gaussian process.

The remainder of the paper is organized as follows. Section 2 gives some background on the theoretical wavelet variance and its estimation for both time series and random fields. Section 3 presents our theory for *M*-estimation of the wavelet variance, with proof of the main result (Theorem 1) deferred to the Appendix. Due to use of a logarithmic transform, our raw *M*-estimators are biased, so Sect. 4 discusses how to correct for this bias. Section 5 considers how to construct CIs for the wavelet variance based upon our *M*-estimators. Sections 6 and 7 discuss computer experiments that investigate the efficacy of our proposed methodology, and Sect. 8 considers a substantive application to two-dimensional cloud data. Section 9 concludes the paper with some discussion.

#### 2 Wavelet analysis of variance

#### 2.1 Daubechies wavelet filter

Let  $\{h_{1,0}, \ldots, h_{1,L-1}\}$  be a unit level Daubechies wavelet filter (Daubechies 1992, Sect. 6.2) of width  $L = L_1$ , which by definition satisfies three conditions:

$$\sum h_{1,l}^2 = \frac{1}{2}, \quad \sum h_{1,l}h_{1,l+2n} = 0,$$

for all non-zero integers *n*, where  $h_{1,l} = 0$  for l < 0 and  $l \ge L$ ; and  $\sum i^l h_{1,l} = 0$  for i = 1, ..., L/2. Let  $H_1(f)$  denote the transfer function (Fourier transform) of the filter  $\{h_{1,l}\}$ . The *j*th level wavelet filter  $\{h_{j,l}\}$  is defined as the inverse discrete Fourier transform (DFT) of

$$H_j(f) = H_1(2^{j-1}f) \prod_{l=0}^{j-2} e^{i2\pi 2^l f(L-1)} H_1(\frac{1}{2} - 2^l f).$$
(1)

The width of this filter is given by  $L_j = (2^j - 1)(L - 1) + 1$ .

## 2.2 Wavelet variance for time series

Let  $\{X_t, t \in \mathcal{Z}\}$  be an intrinsically stationary process of order *d*, where *d* is a non-negative integer; i.e., its *d*th order increments  $(1-\mathcal{B})^d X_t$  are stationary, where  $\mathcal{B}X_t = X_{t-1}$ . Let  $S_X$  denote the spectral density function (SDF) of the process. The *j*th level wavelet coefficient process is then given by

$$W_{j,t} = \sum_{l=0}^{L_j-1} h_{j,l} X_{t-l},$$

which corresponds to the changes on scale  $\tau_j = 2^{j-1}$ . The *j*th level wavelet variance is defined as

$$\nu_j^2 = \operatorname{var}\{W_{j,t}\}.$$

When L = 2, we have the Haar wavelet variance, for which the wavelet coefficients at level *j* are proportional to the difference of simple averages of  $2^{j-1}$  consecutive observations. For L > 2, the wavelet coefficients are contrasts between localized weighted averages. When  $X_t$  is a stationary process with SDF  $S_X$ , Percival (1995) obtained the wavelet variance decomposition

$$\operatorname{var}\left\{X_{t}\right\} = \sum_{j=1}^{\infty} \nu_{j}^{2} \tag{2}$$

as an alternative to the classical decomposition

$$\operatorname{var} \{X_t\} = \int_{-1/2}^{1/2} S_X(f) \, \mathrm{d} f.$$

The decomposition of var  $\{X_t\}$  offered by the wavelet variance complements that of the SDF by focusing directly on scale-based variations, which are often more interpretable and of more interest in the geosciences than frequency-based variations.

Given an observed time series that can be regarded as a realization of  $X_0, \ldots, X_{N-1}$ and assuming the sufficient condition L > 2d to ensure that  $\{W_{j,t}\}$  has zero mean, the usual unbiased estimator of  $v_i^2$  is given by

$$\hat{\nu}_j^2 = \frac{1}{M_j} \sum_{t=L_j-1}^{N-1} W_{j,t}^2, \tag{3}$$

where  $M_j = N - L_j + 1 > 0$ . See Percival (1995) and Percival and Walden (2000) for large sample properties of this estimator and construction of CIs.

## 2.3 Wavelet variance for random fields

Let  $X_{u,v}$ ,  $u, v = 0, \pm 1, \pm 2, ...$  be a stationary Gaussian random field on the twodimensional integer lattice  $\mathbb{Z}^2$  with SDF  $S_X(f, f')$ . Its wavelet transform is defined by filtering the random field using the four possible combinations of wavelet and scaling filters along its rows and columns, yielding four types of coefficients, of which the so-called wavelet–wavelet coefficients are of primary interest to us:

$$W_{j,j',u,v} = \sum_{l=0}^{L_j-1} \sum_{l'=0}^{L_{j'}-1} h_{j,l} h_{j',l'} X_{u-l,v-l'}.$$

The wavelet variance is defined in terms the variance of these coefficients:

$$\nu_{j,j'}^2 = \operatorname{var} \{ W_{j,j',u,v} \}.$$

If  $X_{u,v}$  is intrinsically stationary of order d, then  $S_X(f, f')$  has a pole of order d at the origin and  $v_{j,j'}^2$  is well defined if  $L \ge d$ . The wavelet variance decomposes the process variance since

$$\operatorname{var} \{X_{u,v}\} = \sum_{j=1}^{\infty} \sum_{j'=1}^{\infty} \nu_{j,j'}^2.$$
(4)

This generalizes the result for time series stated in (2) and provides a scale-based analysis of variance for random fields. When we have a realization of an intrinsically stationary random field  $X_{u,v}$  on a finite array {(u, v) : u = 0, ..., N-1, v = 0, ..., M-1}, we can then estimate the wavelet variance by the unbiased estimator

$$\hat{\nu}_{j,j'}^2 = \frac{1}{N_j M_{j'}} \sum_{u=L_j-1}^{N-1} \sum_{v=L_{j'}-1}^{M-1} W_{j,j',u,v}^2, \tag{5}$$

where  $N_j = N - L_j + 1$  and  $M_{j'} = M - L_{j'} + 1$ . See Mondal and Percival (2009) for statistical inference based on this type of estimator.

## 3 M-estimation of wavelet variance

Let  $\{Y_i\}, i \in \mathcal{L}$ , be a zero-mean Gaussian process, and suppose we are interested in estimating var $\{Y_i\} = EY_i^2 = v^2$ . Here the index *i* represents either time *t* or spatial location (u, v), and hence  $\mathcal{L}$  is an integer lattice, either  $\mathcal{Z}$  or  $\mathcal{Z}^2$ . Typically  $\{Y_i\}$  is either the wavelet coefficient process  $\{W_{j,t}\}$  or the wavelet–wavelet coefficients  $\{W_{j,j',u,v}\}$ for any fixed scale (or scale pair). In practice the simple mean-type estimators (3) and (5) are vulnerable to data contamination, so we are interested in developing robust estimators that guard against such contamination, yet still work well when Gaussianity holds. Under our assumptions,  $v^2$  is a scale parameter. A logarithmic transformation converts it to a location parameter and allows use of *M*-estimation theory to formulate a robust estimator. Accordingly, let

$$Q_i = \log Y_i^2$$
.

Then  $\{Q_i\}$  is a stationary process and, using Bartlett and Kendall (1946), we obtain

$$\operatorname{E} Q_i = \log v^2 + \psi(\frac{1}{2}) + \log 2, \quad \operatorname{var} \{Q_i\} = \psi'(\frac{1}{2}) = \frac{\pi^2}{2},$$

where  $\psi$  and  $\psi'$  are the di- and tri-gamma functions. Let  $\mu = \log \nu^2 + \psi(\frac{1}{2}) + \log 2$ . Then  $Q_i$  can be written as

$$Q_i = \mu + \epsilon_i,$$

where  $E\epsilon_i = 0$  and  $var \{\epsilon_i\} = \pi^2/2$ .

**Assumption 1** Let  $\varphi(x), x \in \mathcal{R}$ , be a non-decreasing real-valued function of bounded variation with  $\varphi(-\infty) < 0$  and  $\varphi(\infty) > 0$  such that

$$\lambda(x) = \mathbf{E}\varphi(Q_i - x)$$

is well defined, strictly decreasing on  $\mathcal{R}$  and has a solution point  $\mu_0$  such that

$$\lambda(\mu_0) = 0.$$

Moreover we assume  $\varphi$  is such that  $\lambda(x)$  is differentiable and  $\lambda'(x)$  is continuous around  $\mu_0$ .

The relationship between the solution point  $\mu_0$  and the location parameter  $\mu$  is discussed in Sect. 4.

Because of the Gaussian assumption on  $\{Y_i\}$ , the marginal distribution  $F_Q$  of  $Q_i$  is that of the logarithm of a squared Gaussian random variable and hence is infinitely differentiable. Integration by parts allows us to write

$$\lambda(x) = -\int_{-\infty}^{\infty} F_Q(x+y) \,\mathrm{d}\varphi(y),$$

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and the first derivative of  $\lambda$  satisfies the relation

$$\lambda'(x) = -\int f_Q(y+x) \,\mathrm{d}\varphi(y).$$

Given the form of  $F_Q$  and the fact that  $\varphi$  is of bounded variation, it follows that  $\lambda'(x)$  is bounded as well.

The corresponding *M*-estimator  $T_N$  of the solution point  $\mu_0$  based on observations  $\{Q_i, i \in \mathcal{I}\}$  is defined by

$$T_N = \operatorname{argmin} \left\{ \left| \sum_{i \in \mathcal{I}} \varphi(Q_i - x) \right| : x \in \mathcal{R} \right\}.$$

The index set  $\mathcal{I}$  is equal to  $\{0, ..., N - 1\}$  for time series and is  $\{(u, v) : u, v = 0, ..., N - 1\}$  for a random field (thus, when  $Y_i$  represents  $W_{j,t}$ , N stands for  $M_j$ ). In what follows, let B be the size of  $\mathcal{I}$ .

Assumption 1 holds for various choices of  $\varphi$ , including

$$\varphi_I(x) = \text{sign}(x), \ \varphi_{II}(x) = 2\Pr(\epsilon_i \le x) - 1, \ \varphi_{III}(x) = p \text{ sign}(x) \mathbf{1}_{|x| > p} + x \mathbf{1}_{|x| \le p}$$

for p > 0 and

$$\varphi_{IV}(x) = \begin{cases} a' & \text{if } x \le a, \\ e^x - 1 & \text{if } a < x \le b \\ b' & \text{if } x > b \end{cases} \text{ and }$$

(see Eq. (3.9) of Thall 1979, for another choice). To better understand *M*-estimation, consider an independent and identically distributed (IID) sample. Then  $T_N$  corresponding to  $\varphi_I$  is the same as the maximum likelihood estimator (MLE) for the location parameter when the observations arise from a double exponential distribution. The function  $\varphi_{II}$  corresponds to an MLE under logistic errors, whereas  $\varphi_{III}$  corresponds to an MLE under logistic errors, whereas  $\varphi_{III}$  corresponds to an MLE under a distribution whose central part behaves like a Gaussian but whose tail is like a double exponential. Similarly, for  $\varphi_{IV}$ , the estimator is an MLE under a distribution whose central part behaves like the log of a chi-square distribution. The choice  $\varphi_I$  gives rise to median-type estimators, which, when compared to mean-type estimators, are less sensitive to data contamination. The choice  $\varphi_{II}$  yields Huber's  $\varphi$  function for a location parameter, which maps extreme values of log  $Y_i^2$  to either  $\pm p$ . Similarly  $\varphi_{IV}$  is a Huberized mean of  $Y_i^2$ , which replaces extreme values of  $Y_i^2$  with either a' or b'. The median-type estimator  $\varphi_I$  is invariant to monotone transformation of the data and can be regarded as limiting cases of the Huber-type estimators  $\varphi_{III}$  and  $\varphi_{IV}$ .

*M*-estimation under a non-IID set up has been considered by a large number of authors (see, for example, Beran 1991; Koul and Surgailis 1997). Here we follow the work of Koul and Surgailis (1997), which allows a very general class of weight

functions  $\varphi$ . The following central limit theorem (proven in the Appendix) provides the basis for inference on the solution point  $\mu_0$  using the estimator  $T_N$ .

**Theorem 1** Assume  $\varphi$  and  $\lambda$  satisfy Assumption 1, and  $\{Y_i\}$  has a square integrable SDF. Then  $B^{\frac{1}{2}}(T_N - \mu_0)$  is asymptotically normal with mean zero and variance given by  $A_{\varphi}/[\lambda'(\mu_0)]^2$ , where

$$A_{\varphi} = \sum_{i \in \mathcal{L}} \operatorname{cov}\{\varphi(Q_i - \mu_0), \varphi(Q_0 - \mu_0)\}.$$

Theorem 1 is linked with influence functions and von Mises expansions. When  $\varphi$  is smooth (twice continuously differentiable), for example,  $\varphi = \varphi_{II}$ , then  $\sum_{i \in \mathcal{I}} \varphi(Q_i - T_N) = 0$ . We can then use a Taylor series expansion to deduce that

$$B^{\frac{1}{2}}(T_N - \mu_0) = \frac{B^{-\frac{1}{2}} \sum \varphi(Q_i - \mu_0)}{B^{-1} \sum \varphi'(Q_i - \mu_0) + (T_N - \mu_0)B^{-1} \sum \varphi''(Q_i - T^*)}$$

where  $T^*$  takes values between  $\mu_0$  and  $T_N$ . Consequently the central limit theorem follows from that of  $B^{-\frac{1}{2}} \sum \varphi(Q_i - \mu_0)$  and by proving the consistency of  $T_N$ . However, when  $\varphi$  is no longer smooth, e.g.,  $\varphi = \varphi_I$ , we cannot make use of a Taylor series expansion, so the general proof of Theorem 1 in the Appendix takes a substantially different approach.

#### 4 Correction of bias

The statistics  $T_N$  is consistent for the solution point  $\mu_0$ , which is not necessarily the same as the location parameter  $\mu$ . We can obtain a robust estimator  $\hat{\mu}$  of  $\mu$  by adding bias  $= \mu - \mu_0$  to the estimator  $T_N$ , yielding

$$\hat{\mu} = T_N + \text{bias}.$$

This bias depends on the choice of the weight function  $\varphi$  and on the distribution function  $F_Q$ . We can compute it analytically in some cases. To do so, we first determine the function  $\lambda(x)$ . Let Z,  $\phi$  and  $\Phi$  denote the standard Gaussian random variable and its density and distribution functions. Then

$$\lambda(x) = \mathbf{E}\varphi(Q_i - x) = \mathbf{E}\varphi(\log Z^2 + \log v^2 - x).$$

Thus the choice of  $\varphi = \varphi_I$  yields

$$\lambda_I(x) = 3 - 4\Phi\left(e^{\frac{x}{2} - \log\nu}\right). \tag{6}$$

Therefore,  $\lambda_I(\mu_0) = 0$  implies

$$\mu_0 = 2\log\nu + 2\log\left[\Phi^{-1}(\frac{3}{4})\right].$$
(7)

and hence

bias<sub>I</sub> = 
$$\mu - \mu_0 = \psi(\frac{1}{2}) + \log 2 - 2\log\left[\Phi^{-1}(\frac{3}{4})\right].$$
 (8)

Next consider  $\varphi = \varphi_{II}$ . First we simplify  $\varphi_{II}$  as:

$$\varphi_{II}(x) = 4\Phi\left(e^{\frac{x}{2} + \frac{1}{2}\psi(\frac{1}{2}) + \frac{1}{2}\log 2}\right) - 3.$$

Therefore we obtain

$$\lambda_{II}(x) = 4C_a \left( e^{\log \nu - \frac{x}{2} + \frac{1}{2}\psi(\frac{1}{2}) + \frac{1}{2}\log 2} \right) - 3, \tag{9}$$

where  $C_a$  is the distribution function of a standard Cauchy random variable. Now  $\lambda_{II}(\mu_0) = 0$  implies

$$\mu_0 = 2\log\nu + \psi(\frac{1}{2}) + \log 2 - 2\log\left[C_a^{-1}(\frac{3}{4})\right].$$
 (10)

and hence

bias<sub>II</sub> = 
$$\mu - \mu_0 = 2 \log \left[ C_a^{-1} (\frac{3}{4}) \right]$$
.

For  $\varphi_{III}$  and  $\varphi_{IV}$ , there are no easy closed forms for  $\lambda(x)$ ; however, we can numerically evaluate the bias correction in these cases.

## 5 Construction of confidence intervals

Given a consistent estimator of  $A_{\varphi}$  and knowledge of  $\lambda'(\mu_0)$ , we can use Theorem 1 to construct an asymptotically correct CI for  $\mu_0$  and hence for  $\mu$  and  $\nu^2$ . Since  $A_{\varphi}$  is equal to the SDF of the stationary process { $\varphi(Q_i - \mu_0)$ } at zero frequency, we use a multitaper spectral approach to estimate it (Serroukh et al. 2000). Let { $\gamma_{c,t}$ , t = 0, ..., N-1} for c = 0, ..., C - 1 be the first *C* orthogonal Slepian tapers of length *N* and design bandwidth parameter W = 4/N, where *C* is an odd integer. When dealing with a time series, let  $\mathcal{K}$  be the index set {0, ..., C - 1}; otherwise, for a random field, let it be {(c, c') : c, c' = 0, ..., C - 1}. For  $k \in \mathcal{K}$ , we define

$$J_k = \sum_i \beta_{k,i} \varphi(Q_i - T_N), \qquad (11)$$

where either  $\beta_{k,i} = \gamma_{c,t}$  with i = t for a time series or  $\beta_{k,i} = \gamma_{c,u}\gamma_{c',v}$  with i = (u, v) if we have a random field. Define

$$\tilde{\mu} = \frac{\sum_{k} J_k \beta_{k,.}}{\sum_{k} \beta_{k,.}^2}, \text{ where } \beta_{k,.} = \sum_{i} \beta_{k,i}.$$

We then estimate  $A_{\varphi}$  by

$$\hat{A}_{\varphi} = \frac{1}{K} \sum_{k} (J_{k} - \tilde{\mu} \beta_{k,.})^{2}, \qquad (12)$$

where *K* is the size of the index set  $\mathcal{K}$ . Since  $\mu_0$  is unknown, we use the consistent estimator  $T_N$  in its stead in Eq. (11). Thus the resulting multitaper estimate  $\hat{A}_{\varphi}$  is consistent for  $A_{\varphi}$  if the SDF of the process { $\varphi(Q_i - x)$ } at zero frequency is continuous at  $x = \mu_0$ . The latter holds for a wide range of Gaussian processes { $Y_i$ } and for many choices of  $\varphi$ . In particular, a sufficient condition is that the process { $Y_i$ } is ergodic, for which  $T_N$  is also strongly consistent. Following the recommendation of Serroukh et al. (2000), we choose C = 5 so that the bandwidth of the resulting multitaper estimator is approximately 7/N.

What remains is to compute  $\lambda'(\mu_0)$ . For  $\varphi = \varphi_I$ , we use Eqs. (6) and (7) to obtain

$$\lambda_I'(\mu_0) = -2\phi\left(e^{\frac{\mu_0}{2} - \log\nu}\right)e^{\frac{\mu_0}{2} - \log\nu} = -2\phi\left(\Phi^{-1}(\frac{3}{4})\right)\Phi^{-1}(\frac{3}{4}).$$
 (13)

Similarly for  $\varphi = \varphi_{II}$ , use of Eq. (9) gives

$$\lambda'_{II}(\mu_0) = -2 c_a \left( e^{\log \nu - \frac{\mu_0}{2} + \frac{1}{2}\psi(\frac{1}{2}) + \frac{\log 2}{2}} \right) e^{\log \nu - \frac{\mu_0}{2} + \frac{1}{2}\psi(\frac{1}{2}) + \frac{\log 2}{2}}.$$

where  $c_a$  is the density function of  $C_a$ . By using Eq. (10), we can simplify the above to

$$\lambda'_{II}(\mu_0) = -2 \,c_a \left( C_a^{-1}(\frac{3}{4}) \right) C_a^{-1}(\frac{3}{4}). \tag{14}$$

This has the same form as that of Eq. (13) with the Gaussian density function being replaced by the Cauchy. For other choices of  $\varphi$ , namely,  $\varphi = \varphi_{III}$  and  $\varphi_{IV}$ , there is no convenient analytic form (although one might surmise that it will have a form similar to Eqs. (13) and (14)), but we can evaluate  $\lambda'(\mu_0)$  numerically.

#### 6 Efficiency study

Robust estimators guard against data contamination but are less efficient than estimators designed to be efficient when underlying assumptions are correct. If a time series or random field is truly Gaussian, a robust estimator can perform poorly compared to the mean-type estimator. It is therefore of interest to study the asymptotic relative efficiency (ARE) of the two estimators for a range of Gaussian processes. Theorem 1 yields, approximately,

$$e^{T_N} \sim \text{Log-normal } \left(\mu_0, \frac{\sigma^2}{B}\right), \text{ where } \sigma^2 = \frac{A_{\varphi}}{[\lambda'(\mu_0)]^2}.$$

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Since E exp $(T_N) \approx \exp(\mu_0 + \sigma^2/(2B))$ , an approximately unbiased and robust estimator of  $\nu^2$  is given by

$$\tilde{\nu}^2 = \exp\left(T_N + \operatorname{bias} - \psi\left(\frac{1}{2}\right) - \log 2 - \frac{\hat{\sigma}^2}{2B}\right), \quad \text{where } \hat{\sigma}^2 = \frac{\hat{A}_{\varphi}}{[\lambda'(\mu_0)]^2}.$$
 (15)

Assuming that  $\hat{A}_{\varphi}$  is a consistent estimator of  $A_{\varphi}$ , the ARE is equal to

$$ARE = \lim_{B \to \infty} \frac{\operatorname{var}\{\hat{\nu}^2\}}{\operatorname{var}\{\tilde{\nu}^2\}} = \frac{A[\lambda'(\mu_0)]^2}{\nu^4 A_{\varphi}},$$
(16)

where  $\hat{v}^2 = B^{-1} \sum_i Y_i^2$  is the usual mean-type estimator, which is asymptotically normal with mean  $v^2$  and variance

$$\frac{A}{B} = \frac{2}{B} \sum_{i=-\infty}^{\infty} (\operatorname{cov}\{Y_0, Y_i\})^2$$

(Percival 1995). Using a Hermite expansion, we could write  $A_{\varphi}$  in terms of the autocovariance sequence of  $\{Y_i\}$ , but this expansion is not useful in practice for computing exact AREs. We therefore resort to some simulation studies, specializing to the case  $\varphi = \varphi_I$ , i.e., the median-type estimator.

In the first example, we simulate 10,000 AR(1) time series of length 1,024 for various values of the unit lag correlation  $\phi$  (Kay 1981). For each series we take  $Y_t = W_{2,t}$ , the level j = 2 Haar wavelet coefficients. We compute exact values of  $v^2$ ,  $\mu_0$  and  $\lambda'(\mu_0)$ , so we only need  $A/A_{\varphi}$  to determine the ARE via (16). This ratio can be approximated by var { $\sum Y_t^2$ }/var { $\sum \varphi_I (\log Y_t^2 - \mu_0)$ }. Hence for each time series we compute  $\sum Y_t$  and  $\sum \varphi_I (\log Y_t^2 - \mu_0)$  and then compute their corresponding sample variances using all 10,000 replications. Figure 3a plots the estimated ARE against  $\phi \in (-1, 1)$ . The ARE is about 50% for all  $\phi$ , is smallest when  $\phi$  is close to -1 and attains its peak value at about  $\phi = 0.75$ , above which it then declines slightly. As an additional check of our theory, we note that we obtain similar results if we just compute the ratio of the Monte Carlo variances for  $\hat{v}^2$  and  $\tilde{v}^2$  as suggested by Eq. (16).

For the second example, we consider a stationary fractionally differenced (FD) process with long memory parameter  $\delta \in (0, \frac{1}{2})$  ( $\delta = 0$  corresponding to white noise, and the process becomes more highly correlated as  $\delta$  approaches 1/2). For selected  $\delta$  we simulate 10,000 FD time series of length 1,024 (Craigmile 2003) and estimate the ARE for median-type versus mean-type estimators as in the AR(1) example. Figure 3b plots the ARE versus  $\delta$ . We see that the ARE is close to 50% for all  $\delta$ .

A simulation experiment with fractional Brownian surfaces (see, e.g., Zhu and Stein 2002) gives about 60–65% efficiency (details are omitted).



Fig. 3 Approximate asymptotic relative efficiency of the median-type estimator with respect to the meantype estimator for **a** AR(1) processes with parameter  $\phi$  and **b** FD processes with parameter  $\delta$ 

## 7 Simulation study

Here we report the results of a simulation study that demonstrates the efficacy of our theory and that expands upon the example considered in Figs. 1 and 2. We consider a Gaussian stationary process that models a burst-free portion of an actual ocean shear series previously considered in Percival (1995) and Percival and Walden (2000). The SDF for this model is piecewise power law (Percival and Walden 2000, p. 331). We generate realizations of length N = 4,096 using the Gaussian spectral synthesis method (Percival and Walden 2000, Sect. 7.8; the parameter M described there is set to 4N). We compute the usual (mean-type) unbiased estimator  $\hat{\nu}_i^2$  of the wavelet variance for scales  $\tau_i = 2^{j-1}$ ,  $j = 1, \dots, 9$ , using Eq. (3) with  $\{h_{j,l}\}$  based upon the D(6) wavelet filter, i.e., the Daubechies extremal phase filter of width L = 6 (Daubechies 1992, Table 6.1). We then compute an approximately unbiased median-type robust estimator  $\tilde{\nu}_i^2$  using  $\varphi = \varphi_I$ . To do so, we let  $T_N$  be the log of the median of the same squared coefficients used to form  $\hat{v}_i^2$  and then use it in Eq. (15), along with the bias term given by Eq. (8),  $\hat{A}_{\varphi}$  by (12) and  $\lambda'(\mu_0)$  by (13). We next corrupt the simulated time series by taking its orthonormal D(6) discrete wavelet transform (DWT) and multiplying selected level j = 3 coefficients by log Gaussian white noise exp $(\epsilon_t)$  with  $E\epsilon_t = 0$  and  $var\{\epsilon_t\} = 1.5$ . There are 512 DWT coefficients at level j = 3, of which we select 51 at random for alteration. Additionally, we alter randomly chosen patches of coefficients, where the patchiness is dictated by a realization of a stationary Markov chain  $\eta_t$  with

$$Pr(\eta_t = 1 | \eta_{t-1} = 0) = 0.09, Pr(\eta_t = 0 | \eta_{t-1} = 1) = 0.01$$

and  $Pr(\eta_t = 0) = 0.1$ . Any level j = 3 coefficient with a corresponding  $\eta_t$  of zero is multiplied by log Gaussian white noise with the same statistical properties as before. The total number of altered coefficients on the average is approximately 97 allowing some coefficients to be altered twice. We take the inverse DWT to create a contaminated version of the original simulated time series [note that, although we have altered just the level j = 3 DWT coefficients, the wavelet coefficients that are used to estimate the wavelet variance are based on the over-complete maximal overlap DWT (Percival and Walden 2000), for which we can expect the contamination to leak out into scales adjacent to  $\tau_3$ ]. We then compute the mean- and median-type wavelet

variance estimators. Finally we repeat this entire process over again for 1,000 different realizations.

Tables 1 and 2 summarize the results of our simulation study. Table 1 concerns just the uncontaminated series. The second column shows the average of the 1,000 estimates of  $\hat{v}_j^2$ . The third column shows the ratio of the average of the median-type estimates  $\tilde{v}_j^2$  to the average of the mean-type estimates  $\hat{v}_j^2$ . With the exception of the largest scale  $\tau_9$ , these ratios are very close to unity, which indicates that the mean- and median-type estimates match up quite well. The poorer agreement at scale  $\tau_9$  can be attributed to a downward bias in the estimator  $\hat{A}_{\varphi}$  due to a sparsity of relevant data at that scale (if we were to replace  $\hat{A}_{\varphi}$  in (15) with its actual value, the modified version of  $\tilde{v}_9^2$  would agree well with  $\hat{v}_9^2$ ). The fourth and fifth columns give the sample standard deviations (SDs) for the 1,000 estimates of, respectively,  $\hat{v}_j^2$  and  $\tilde{v}_j^2$ . The final column gives the estimated relative efficiency of the median-type estimator to the mean-type estimator. It is interesting that the efficiency is markedly smaller for scale  $\tau_1$  than for scales  $\tau_2$  to  $\tau_8$ . The small efficiency for scale  $\tau_9$  can again be attributed to lack of

j	Mean $\{\hat{v}_j^2\}$	$\tilde{v}_j^2/\hat{v}_j^2$	SD $\{\hat{v}_j^2\}$	$\operatorname{SD} \{ \tilde{v}_j^2 \}$	Relative efficiency (%)
1	0.000585	1.00	0.000015	0.000022	48.1
2	0.000904	1.00	0.000031	0.000041	57.7
3	0.001743	1.00	0.000080	0.000102	60.7
4	0.007150	1.00	0.000550	0.000693	63.0
5	0.040047	1.00	0.004489	0.005638	63.4
6	0.214502	1.01	0.034153	0.042587	64.3
7	0.858346	1.01	0.216022	0.269830	64.1
8	1.197886	1.03	0.422244	0.547121	59.6
9	0.641284	1.27	0.367127	0.567690	41.8

Table 1 Summary of wavelet variance estimates of uncontaminated simulated ocean shear data

Table 2 Summary of wavelet variance estimates of contaminated ocean shear data

j	$\hat{v}_{C,j}^2/\hat{v}_j^2$	$\tilde{v}_{C,j}^2/\hat{v}_j^2$	$SD \{\hat{v}_{C,j}^2\}$	SD $\{\tilde{\nu}_{C,j}^2\}$	RMSE $\{\hat{v}_{C,j}^2\}$	RMSE $\{\tilde{v}_{C,j}^2\}$
1	12.35	1.05	0.160084	0.000028	0.160142	0.000042
2	62.68	1.12	1.366784	0.000065	1.367238	0.000130
3	72.02	1.13	3.046282	0.000157	3.047273	0.000278
4	4.98	1.07	0.693477	0.000743	0.693714	0.000885
5	1.04	1.01	0.035961	0.005404	0.035981	0.005413
6	1.01	1.00	0.033752	0.043243	0.033762	0.043225
7	1.00	1.01	0.220646	0.274201	0.220536	0.274209
8	0.97	0.99	0.385875	0.504024	0.387926	0.503808
9	1.02	1.32	0.391146	0.668527	0.391120	0.698292

sufficient data relevant to this scale (if we increase the sample size to  $N = 4 \times 4,096$ , the efficiency increases back up to above 60%).

Table 2 shows results involving the contaminated series, where  $\hat{v}_{C,j}^2$  and  $\tilde{v}_{C,j}^2$  denote  $\hat{v}_j^2$  and  $\tilde{v}_j^2$  when applied to the contaminated series. The second column shows the ratio of the average of the mean-type estimates  $\hat{v}_{C,j}^2$  for the contaminated series to a similar average for the uncontaminated series, while the third column shows a similar ratio, but now involving the average of the median-type estimates  $\tilde{v}_{C,j}^2$  for the contaminated series. Contamination has a very adverse effect on the mean-type wavelet variance estimates at small scales ( $\tau_1-\tau_4$ ), whereas the median-type estimates are far more resistance to contaminated data. Note that the contamination has little effect on the mean- or median-type estimates at scales  $\tau_5$  and above (the relatively poor performance of  $\tilde{v}_{C,9}^2$  can again be attributed to bias in the estimator  $\hat{A}_{\varphi}$ ). The remaining columns in the table show the sample SDs and root mean-square errors (RMSEs) for the 1,000 estimates of  $\hat{v}_{C,j}^2$  and  $\tilde{v}_{C,j}^2$ . The RMSEs suggest that the robust estimator performs better up to scale  $\tau_4$ , beyond which it is better to switch over to the mean-type estimator.

The time series that we selected from amongst the 1,000 series to show in Fig. 1 is typical in the sense that its level j = 3 actual RMSE for the contaminated mean-type estimate is closest to the sample RMSE for the 1,000 such estimates. The CIs based upon this series which are displayed in Fig. 2 are thus also typical of what we can expect to get. Note that the CIs for  $v_9^2$  based upon the median-type estimates are markedly smaller than those based upon the mean-type estimates, which is not the case for scales  $\tau_1 - \tau_8$ . This anomaly can again be attributed to bias in the estimator  $\hat{A}_{\varphi}$  due to lack of sufficient data at that scale.

#### 8 Application to cloud data

Figure 4 shows the pre-processed brightness temperature image of a cloud field over the southeast Pacific Ocean obtained on 17 October 2001 as part of the East Pacific Investigation of Climate (EPIC) field experiment (Bretherton et al. 2004). Stratocumulus cloud fields in that part of the world tend to be homogenous except for pockets of seemingly cloud-free air. These pockets of open cells (POCs) are distinct from broken clouds, are coupled to the development of marine rainfall and are characterized by low-aerosol air mass (Stevens et al. 2005). The four squares in Fig. 4 indicate regions with four different types of clouds. Region (a) contains POCs; (b) consists of uniform stratus clouds and thus has different characteristics than the POC region; (c) has broken clouds; and (d) has clouds that are forming a POC.

Satellite cloud data are often marked by contamination from various sources, but the one of interest to us here is the presence of aberrant cloud types in a region where a single cloud type is dominant. The four regions of focus in Fig. 4 all have a dominant cloud type, but they are homogeneous to varying degrees, with regions (a) and (b) being visually more so than (c) and (d). It is thus natural to resort to median-type estimators that are robust and effective in extracting the characteristics of the dominant cloud type.



Fig. 4 Image plots of clouds at four different regions



**Fig. 5** Log of wavelet variances at diagonal scales indexed by j = 1, 2, 3 and 4 for four cloud regions. The *gray* and *black squares* show, respectively, the mean-type and robust median-type estimates. The *vertical lines* bisecting the squares depict 95% CIs. **a** For the POC region, **b** for uniform clouds, **c** for broken clouds and **d** for a forming POC region

Figure 5 shows conventional mean-type (gray squares) and robust median-type (black) wavelet variance estimates and their 95% CIs (lines intersecting squares vertically) for four diagonal scales (i.e., j' = j) indexed by j = 1, 2, 3 and 4. Plot (a) is for the POC region. At the three smallest scales ( $\tau_1-\tau_3$ ) the median-type estimates take somewhat lower values than those for the mean-type estimates, with the largest discrepancy occurring at scale  $\tau_1$ . While this pattern is consistent with this region being contaminated to some degree by cloud types other than POCs, the fact that the mean- and median-type estimates are comparable to within the sampling variability indicated by the associated overlapping CIs suggests that the POC region is largely

unaffected by other cloud types. Plot (b) is for the uniform stratus clouds. Again the mean- and median-type wavelet variance estimates and the associated CIs suggest that this region is largely homogenous; however, unlike the POC region, the CI based upon median-type estimate at scale  $\tau_1$  is markedly wider than the one based upon the mean-type estimate, which leads to the speculation that the noise characteristics of the two regions are different. Plots (c) and (d) show the results for the regions with broken clouds and POC formation. At scales  $\tau_1$  and  $\tau_2$ , the robust estimates of wavelet variances for the broken clouds are almost an order of magnitude lower than the conventional estimates, but also have much larger CIs. Broken clouds are mixtures of various clouds, so the median-type estimates should pick out the characteristics from the dominant cloud type; however, a larger fraction of other cloud types produces bigger CIs. The robust estimates of wavelet variance for the region of POC formation are again much smaller than for the conventional estimates, evidently because of the presence of other cloud types; however, unlike the broken clouds region, the CIs based upon the robust procedure are smaller in this region, indicating that the fraction of other clouds present here is smaller than the fraction in the broken clouds region.

One potential use for the wavelet variance in this application is to define features that can be used to classify cloud regions on other images. The wavelet variance curve for a POC region is uniquely identified by a monotonic decrease across scales  $\tau_1 - \tau_4$ , while an overall low level is indicative of uniform clouds. In both cases, the mean- and median-type estimates are comparable. In contrast, these two estimates are markedly different for the broken clouds region and the region of POC formation, with large uncertainty in the median-based estimate being a potential identifier for the broken clouds region. More research is needed to determine if these patterns persist enough across other images to serve as usual features for classification.

## 9 Discussion

The subject of robustness in statistics has been around for more than a quarter of a century, but the usefulness of robust estimators in practice has been the subject of some controversy; e.g., Stigler (1977) questions the usefulness of the median as an estimator of the location parameter for real data. We have seen in Sect. 8 that meanand median-type estimates of the wavelet variance provide different answers at certain scales for inhomogeneous cloud regions. The robust estimates arguably allow us to pick out the characteristics of the dominant cloud type better, thus providing an argument for the usefulness of median-type estimates.

While a number of different contamination models has been entertained for time series data in the literature, we have introduced a new model based upon the idea of scale-based multiplicative contamination. This model is based on the supposition that contamination can occur and affect data at certain scales. Our computer experiments indicate that, even if contamination is present at some small scales, the larger scales will not be influenced much and should have wavelet coefficients that are close to Gaussian. Table 2 shows that the mean-type estimators at larger scales produce smaller mean square error in comparison to robust median-type estimators. A step forward would be to identify the contaminated scales and devise a hybrid scheme

whereby we use the robust estimator for scales at which data are affected by contamination, and then switch over to the mean-type estimator at other scales. This strategy is used in certain wavelet shrinkage problems involving non-Gaussian data (Gao 1997). A question for future research is how to identify the scales at which to switch over in practice.

The wavelet variance provides a simple and useful estimator of the integral of the SDF over octave bands. In particular, the Blackman and Tukey (1958, Sect. 18) pilot spectra coincide with Haar wavelet variances. Recently Tsakiroglou and Walden (2002) extended the pilot spectra of Blackman and Tukey by utilizing the (maximum overlap) discrete wavelet packet transform. The result is an SDF estimator that is competitive with existing estimators. In the same vein, our proposed methodology can be extended to handle wavelet packet transforms, thus providing a robust SDF estimator in the style of Tsakiroglou and Walden (2002).

In practical applications, the Huber function  $\phi = \phi_{IV}$  is of considerable interest because  $\exp(T_N)$  gives rise to the median-type estimator when -a = b = 0 with -a' = b' = 1 and to the mean-type estimator when  $-a = b = \infty$ . With a' and b'set appropriately, other values of -a = b = h can the thought of as a compromise between the robust and the conventional estimation procedures. It is thus of interest to set h so that the Huberized estimator has a certain ARE as given in (16). Following Koul and Surgailis (1997), we could obtain an expression for the ARE via the Hermite expansion and then try to find an h that achieves the required efficiency. This would, however, require knowledge of the autocovariance sequence of the underlying process. A better strategy is to estimate the ARE in (16) by a non-parametric (multitaper) estimator for any given h and then solve an optimization problem on a finite grid for a range of values of h.

Although not considered in the present paper, wavelet-based analysis of variance can be applied to multivariate processes; for example, Whitcher et al. (2000) discuss wavelet covariance analysis of two time series. Wavelet variances and covariances from multiple time series can be used to form a wavelet dispersion matrix, which is useful for studying scale-based correlations and can lead to scale-based versions of clustering, classification and principle component analysis. A robust estimator of this dispersion matrix becomes of interest when multivariate data (time series or random fields) are subject to contamination. Following the discussion of Maronna et al. (2006, p. 205), the methodology we have described can be adapted to form robust pairwise estimators of wavelet variances and covariances. As is true in the univariate case, these robust estimators are asymptotically normal, which provides a basis for drawing inferences; however, for finite sample sizes, these estimators might not yield a dispersion matrix that is non-negative definite (even though asymptotically they will). There are several alternative approaches to devising robust estimators of the wavelet dispersion matrix, including *M*-estimators, *S*-estimators, Stahel–Donoho estimators, minimum covariance determinant estimators, and orthogonalized Gnandesikan-Kettenring estimators (for details, see Maronna 1976; Tyler 1987; Croux and Haesbroeck 2000; chapter 6 of Maronna et al. 2006 and the references therein). A study of the strengths and weaknesses of all these estimators is a subject for future research.

#### **Appendix: Proof of Theorem 1**

We denote by  $\{P_n, n = 0, 1, ...\}$  the sequence of Hermite polynomials. Let  $\phi$  be the density function for a standard Gaussian random variable *Z*. For a function *P* with  $\int_{-\infty}^{\infty} P(x)\phi(x) dx < \infty$ , we write the Hermite expansion as

$$P(x) = \sum_{n=0}^{\infty} c_n P_n(x), \quad \text{where } c_n = \int_{-\infty}^{\infty} P(x) P_n(x) \phi(x) \, \mathrm{d}x. \tag{17}$$

The function P(x) is said to have Hermite rank  $\rho$  if in the expansion (17) we have

$$c_0 = c_1 = \dots = c_{\varrho-1} = 0, \ c_\varrho \neq 0.$$

Prior to proving the theorem, we need to prove the following six lemmas.

Lemma 1 The following functions have Hermite rank 2.

(i) 
$$P(x) = x^2 - 1$$
  
(ii)  $P(x) = \log(x^2) - \int \log(x^2)\phi(x) dx$   
(iii)  $P(x) = 1_{(\log x^2 \le y)} - \int 1_{(\log x^2 \le y)}\phi(x) dx, y \in \mathcal{R}$   
(iv)  $P(x) = \varphi(\log x^2 - y) - \int \varphi(\log x^2 - y)\phi(x) dx, y \in \mathcal{R}$ .

*Proof of Lemma* 1. Recalling that  $P_0(x) = 1$ ,  $P_1(x) = x$  and  $P_2(x) = x^2 - 1$ , note that each P(x) is an even function such that  $\int P(x)\phi(x) dx = 0$ , implying that  $c_0 = c_1 = 0$ , whereas  $\int P(x)(x^2 - 1)\phi(x) dx \neq 0$ .

**Lemma 2** If P is any of the functions in Lemma 1 and if  $\{Y_i\}$  has a square integrable spectral density, then the autocovariances  $s_{P,k}$  of the random process  $\{P(Y_i)\}$  satisfy

$$\sigma_P^2 = \sum_k s_{P,k} > 0. \tag{18}$$

Proof of Lemma 2. Since P is even, we can use the Hermite expansion to write

$$P(Y_i) = \sum c_{2m} P_{2m}(Y_i).$$

The SDF of  $\{P(Y_i)\}$  is given by

$$S_P(f) = \sum c_{2m}^2 (2m)! S_Y^{(*2m)}(f)$$

(see Hannan 1970, p. 83), where  $S_Y^{(*2m)}$  is 2*m*-fold convolution of  $S_Y$ . Since  $S_Y^{(*2m)}$  is strictly positive at the origin and there exists one *m* such that  $c_{2m} \neq 0$ , we see that  $\sigma_P^2 = S_P(0)$  is strictly positive, and hence the result follows.

**Lemma 3** If P is any of the functions in Lemma 1 and if  $\{Y_i\}$  has a square integrable SDF, then, as  $N \to \infty$ ,  $B^{-1} \sum_{i \in \mathcal{I}} P(Y_i)$  converges in distribution to  $\sigma_P Z$ , where  $\sigma_P^2$  is as in Eq. (18).

*Proof of Lemma* 3. This follows directly from Lemmas 1, 2 and Theorem 2 of Breuer and Major (1983).

**Lemma 4** Let  $\Psi_N(x) = B^{-1} \sum \varphi(Q_i - x)$  and  $T_N$  be defined as earlier. Then  $T_N - \mu_0 = o_P(1)$  and  $\Psi_N(T_N) = O_P(B^{-1})$ .

*Proof of Lemma* 4. Let  $F_{Q,N}$  be the empirical distribution function of  $\{Q_i, i \in \mathcal{I}\}$ . As  $Q_i \neq Q_j, i \neq j$  a.s. (almost surely), the jumps of the empirical distribution  $F_{Q,N}$  are such that  $\Delta F_{Q,N}(x) = F_{Q,N}(x) - F_{Q,N}(x-1) \leq B^{-1}$  a.s., and therefore  $\Delta \Psi_N(x) = O(B^{-1})$  a.s.; indeed, a.s.

$$|\Delta \Psi_N(x)| \le \int |\Delta F_{\mathcal{Q},N}(y+x)| |\mathrm{d}\varphi(y)| \le |\varphi|/B,$$

where  $|\varphi|$  is the variation of  $\varphi$ . Now since  $\varphi(-\infty) < 0$  and  $\varphi(\infty) > 0$ , we have that  $\Psi_N(-\infty) > 0$  and  $\Psi_N(\infty) < 0$ . Since  $\Psi_N(x)$  is non-increasing, the graph of  $\Psi_N$  crosses the *x*-axis in a neighborhood of 0 at some point  $T_N$  with  $\Psi_N(T_N+) \le 0$  and  $\Psi_N(T_N-) \ge 0$  and hence  $|\Psi_N(T_N+)| + |\Psi_N(T_N-)| = |\Psi_N(T_N-) - \Psi_N(T_N+)| \le |\varphi|/B$ . Hence, for all *N*, we have

$$|\Psi_N(T_N)| \le |\varphi|/B \quad \text{a.s.} \tag{19}$$

We now prove consistency of  $T_N$ . Let  $\epsilon > 0$ . Since  $\Psi_N(x)$  is non-increasing, we note that

$$\Pr(T_N < \mu_0 + \epsilon) > \Pr(\Psi_N(T_N) > \Psi_N(\mu_0 + \epsilon)).$$

By Eq. (19), it then follows that

$$\Pr(\Psi_N(T_N) > \Psi_N(\mu_0 + \epsilon)) > \Pr(|\varphi|/B > \Psi_N(\mu_0 + \epsilon)).$$

However, by Lemma 3,  $\Psi_N(\mu_0 + \epsilon)$  converges in probability to  $\lambda(\mu_0 + \epsilon)$ , which is strictly less than zero. Thus  $\Pr(|\varphi|/B > \Psi_N(\mu_0 + \epsilon))$  converges to one. Hence

$$\Pr(T_N < \mu_0 + \epsilon) \rightarrow 1.$$

A similar argument implies that  $Pr(T_N > \mu_0 - \epsilon)$  converges to one, from which the consistency of  $T_N$  follows.

**Lemma 5** For  $y \ge 0$  we have

$$BE\left[\int \{\varphi(x-y) - \varphi(x)\} d\{F_{Q,N}(x) - F_Q(x)\}\right]^2 \le \text{ constant } y.$$

*Proof of Lemma* 5. Consider the Hermite expansion of  $P(Y_i) = 1_{\log Y_i^2 \le x} - F_Q(x)$ , namely,

$$1_{\log Y_i^2 \le x} - F_Q(x) = \sum_m a_{2m}(x) P_{2m}(Y_i).$$

Lemma 1 says that P has Hermite rank 2. For  $x \le y$  we define  $F_Q(x, y) = F_Q(y) - F_Q(x)$ ,  $a_{2m}(x, y) = a_{2m}(y) - a_{2m}(x)$  and  $F_{Q,N}(x, y) = F_{Q,N}(y) - F_{Q,N}(y)$ . Then we can write

$$1_{x \le \log Y_i^2 \le y} - F_Q(x, y) = \sum_m a_{2m}(x, y) P_{2m}(Y_i).$$

It then follows from the orthogonality of the Hermite polynomials that

$$\sum a_{2m}^2 (2m)! s_{Y,0}^{2m} = \mathbb{E} \{ \mathbb{1}_{x \le \log Y_i^2 \le y} - F_Q(x, y) \}^2 \le F_Q(x, y).$$
(20)

Now

$$E\left[\sum_{i\in\mathcal{I}} \{1_{x\leq \log Y_i^2\leq y} - F_Q(x, y)\}\right]^2 = \operatorname{var}\left\{\sum_m a_{2m}(x, y) \sum_i P_{2m}(Y_i)\right\}^2$$
$$= \sum_m \sum_{m'} a_{2m}(x, y) a_{2m'}(x, y) \sum_i \sum_{i'} \operatorname{cov}(P_{2m}(Y_i), P_{2m'}(Y_{i'})).$$

However, it follows from Hannan (1970, p. 117) that  $\operatorname{cov}(P_{2m}(Y_i), P_{2m'}(Y_{i'})) = 0$  for  $m \neq m'$  and  $\operatorname{cov}(P_{2m}(Y_i), P_{2m}(Y_{i'})) = s_{Y,i-i'}^{2m}$ . Therefore,

$$\mathbb{E}\left[\sum_{i\in\mathcal{I}}\{1_{x\leq\log Y_{i}^{2}\leq y}-F_{Q}(x,y)\}\right]^{2}=\sum_{m}a_{2m}^{2}(x,y)(2m)!\sum_{i}\sum_{i'}s_{Y,i-i'}^{2m}.$$

Let  $\{r_{Y,k} = s_{Y,k}/s_{Y,0}, k \in \mathcal{L}\}$  be the autocorrelation sequence of  $\{Y_i\}$ . Then, by Eq. (20), we obtain

$$\sum a_{2m}^2(x, y)(2m)! \sum \sum s_{Y, i-i'}^{2m} \le F_Q(x, y) \sum_i \sum_{i'} r_{Y, i-i'}^2.$$

Hence

$$BE\left\{F_{Q,N}(x, x+y) - F_{Q}(x, x+y)\right\}^{2} = B^{-1}E\left[\sum_{i \in \mathcal{I}} \{1_{x \le \log Y_{i}^{2} \le y} - F_{Q}(x, y)\}\right]^{2}$$
  
$$\leq F_{Q}(x, y)B^{-1}\sum_{i}\sum_{i'} r_{Y,i-i'}^{2} \le \text{ constant } F_{Q}(x, y).$$
(21)

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The last inequality follows since the SDF of  $\{Y_i\}$  is square integrable. Next we note that

$$\begin{split} &\int \{\varphi(x-y) - \varphi(x)\} \, \mathrm{d}\{F_{Q,N}(x) - F_Q(x)\} \\ &= \int \left\{F_{Q,N}(x+y) - F_{Q,N}(x) - F_Q(x+y) + F_Q(x)\right\} \, \mathrm{d}\varphi(x) \\ &= \int \{F_{Q,N}(x,x+y) - F_Q(x,x+y)\} \, \mathrm{d}\varphi(x), \end{split}$$

so that

$$\left[\int \{\varphi(x-y) - \varphi(x)\} d\{F_{Q,N}(x) - F_Q(x)\}\right]^2$$
  
$$\leq \int \{F_{Q,N}(x, x+y) - F_Q(x, x+y)\}^2 d\varphi(x).$$

Taking expectation we obtain

$$\mathbb{E}\left[\int \{\varphi(x-y) - \varphi(x)\} d\{F_{Q,N}(x) - F_Q(x)\}\right]^2$$
  
 
$$\leq \int \mathbb{E}\left\{F_{Q,N}(x, x+y) - F_Q(x, x+y)\right\}^2 d\varphi(x).$$

So Eq. (21) yields

$$BE\left[\int \{\varphi(x-y) - \varphi(x)\} d\{F_{Q,N}(x) - F_Q(x)\}\right]^2$$
  

$$\leq \int BE\left\{F_{Q,N}(x, x+y) - F_Q(x, x+y)\right\}^2 d\varphi(x)$$
  

$$\leq \text{constant} \int F_Q(x, x+y) d\varphi(x)$$
  

$$\leq \text{constant} \int y \sup f_Q(z) d\varphi(x) \leq \text{constant } y,$$

where the last inequality follow since the density function  $f_Q(x)$  of  $F_Q(x)$  is bounded. This completes the proof.

# Lemma 6

$$h_N = B^{\frac{1}{2}} \int \{\varphi(x - T_N) - \varphi(x - \mu_0)\} d\{F_{Q,N}(x) - F_Q(x)\} = o_P(1).$$

*Proof of Lemma* 6. Assume WLOG  $\mu_0 = 0$ . Then we note that

$$\Pr(|h_N| > \delta) \le \Pr\left[\sup_{|y| < B^{-\gamma}} B^{\frac{1}{2}} | \int \{\varphi(x-y) - \varphi(x)\} d\{F_{Q,N}(x) - F_Q(x)\}| > \delta\right]$$
$$+ \Pr(|T_N| > B^{-\gamma}).$$

The second term is  $o_P(1)$  by Lemma 4. The first term follows by mimicking the chaining argument in Lemma 2.2 of Koul and Surgailis (1997). As in Koul and Surgailis (1997), we prove the result for  $0 \le y \le B^{-\gamma}$  and  $\varphi$  non-decreasing. We put  $y_B = B^{-\gamma}$ and let

$$K = \lfloor \log_2(By_B) \rfloor.$$

We consider a sequence of partitions

$$\{x_{i,k} = y_B i 2^{-k}, 0 \le i \le 2^k\}, k = 0, 1, \dots, K$$

of intervals  $[0, y_B]$ . For a y in  $[0, y_B]$  and a k in  $\{0, 1, \dots, K\}$ , we define i(k, y) by

$$x_{i(k,y),k} \le y < x_{i(k,y)+1,k}.$$

We then obtain a chain by linking 0 to a given point  $y \in [0, y_B]$  as

$$0 = x_{i(0,y),0} \le x_{i(1,y),1} \le \cdots \le x_{i(K,y),K} \le y < x_{i(K,y)+1,K}.$$

Let

$$R_N(y) = B^{\frac{1}{2}} \int \{\varphi(x-y)\} d\{F_{Q,N}(x) - F_Q(x)\}$$
  
=  $B^{\frac{1}{2}} \int \{F_{Q,N}(x+y) - F_Q(x+y)\} d\varphi(x),$ 

and  $R_N(y, z) = R_N(z) - R_N(y)$ . We can then use the above chain to write

$$R_N(0, y) = R_N(x_{i(0,y),0}, x_{i(1,y),1}) + R_N(x_{i(1,y),1}, x_{i(2,y),2}) + \dots + R_N(x_{i(K-1,y),K-1}, x_{i(K,y),K}) + R_N(x_{i(K,y),K}, y).$$

Hence

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$$\sup_{y \in [0, y_B]} R_N^2(0, y) \\ \leq 2 \left( \sum_{k=0}^{K-1} \sup_{y \in [0, y_B]} |R_N(x_{i(k-1, y), k-1}, x_{i(k, y), k})| \right)^2 + 2 \sup_{y \in [0, y_B]} R_N^2(x_{i(K, y), K}, y).$$

We now apply Cauchy-Schwartz inequality to obtain

$$\mathbb{E} \sup_{y \in [0, y_B]} R_N^2(0, y) \le 2K \sum_{k=0}^{K-1} \mathbb{E} \sup_{y \in [0, y_B]} R_N^2(x_{i(k-1, y), k-1}, x_{i(k, y), k})$$
  
+ 2\mathbf{E} \sup\_{y \in [0, y\_B]} R\_N^2(x\_{i(K, y), K}, y).

We now give a bound to the last term. We use the monotonicity of  $F_{Q,N}$ , boundedness of  $\varphi$ , and the fact that  $F_Q$  is the distribution of log of a chi-square random variable. We obtain

$$|R_N(x_{i(K,y),K}, y)| = B^{\frac{1}{2}} \left| \int F_{Q,N}(z + x_{i(K,y),K}, z + y) \, \mathrm{d}\varphi(z) - \int F_Q(z + x_{i(K,y),K}, z + y) \, \mathrm{d}\varphi(z) \right|,$$

which is less than or equal to

$$B^{\frac{1}{2}} \int F_{Q,N}(z + x_{i(K,y),K}, z + x_{i(K,y)+1,K}) d\varphi(z) + \text{ constant } B^{\frac{1}{2}} y_B 2^{-K}.$$

The above is also less than or equal to

$$|R_N(x_{i(K,y),K}, x_{i(K,y)+1,K})| + \text{ constant } B^{\frac{1}{2}}y_B 2^{-K}$$

for a different choice of constant.

Next we observe that for  $k = 0, 1, \ldots, K - 1$ 

$$\sup_{y \in [0, y_B]} |R_N(x_{i(k, y), k}, x_{i(k, y)+1, k+1})|$$
  
= 
$$\max_{0 \le i \le 2^{k+1} - 1} \sup_{y \in [x_{j, k+1}, x_{j+1, k+1}]} |R_N(x_{i(k, y), k}, x_{i(k, y)+1, k+1})|$$
  
$$\le \max_{0 \le i \le 2^{k+1} - 1} |R_N(x_{i, k+1}, x_{i+1, k+1})|$$

Hence in view of Lemma 5, we get

$$E \sup_{y \in [0, y_B]} R_N^2(x_{i(k, y), k}, x_{i(k, y)+1, k+1})$$
  
$$\leq \sum_{i=0}^{2^{k+1}-1} E R_N^2(x_{i, k+1}, x_{i+1, k+1}) \leq \text{constant } y_B,$$

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and similarly

$$\mathbb{E} R_N^2(x_{i(K,y),K}, x_{i(K,y)+1,K}) \le \sum_{i=0}^{2^K-1} \mathbb{E} R_N^2(x_{i,K}, x_{i+1,K}) \le \text{ constant } y_B.$$

Consequently,

$$\operatorname{E}\sup_{y\in[0,y_B]} R_N^2(0,y) \le \operatorname{constant} y_B K^2 + \operatorname{constant} B y_B^2 2^{-2K}$$

Now from the definition of *K*, we obtain  $2^{-2K} = B^{-2(1-\gamma)}$ , and thus

$$By_B 2^{-2K} = O(B^{1-2\gamma-2+2\gamma}) = O(B^{-1}), \quad K^2 y_B = O(B^{-\gamma} \log_2^2 B).$$

This completes the proof.

Proof of Theorem 1. By virtue of Lemma 4, we can write

$$O_P(B^{-1}) = \Psi_N(T_N) = B^{-1} \sum_{i \in \mathcal{I}} \varphi(Q_i - T_N) - \int \varphi(x - \mu_0) \, \mathrm{d}F_Q(x)$$
  
$$= \int \varphi(x - T_N) \, \mathrm{d}F_{Q,N}(x) - \int \varphi(x - \mu_0) \, \mathrm{d}F_Q(x)$$
  
$$= \int \varphi(x - T_N) \, \mathrm{d}\{F_{Q,N}(x) - F_Q(x)\}$$
  
$$+ \int \{\varphi(x - T_N) - \varphi(x - \mu_0)\} \, \mathrm{d}F_Q(x)$$
  
$$= \rho_N + \lambda(T_N) - \lambda(\mu_0),$$

where  $\rho_N = \int \varphi(x - T_N) d\{F_{Q,N}(x) - F_Q(x)\}$ . This implies

$$\lambda(T_N) - \lambda(\mu_0) = o_P(B^{-\frac{1}{2}}) - \rho_N.$$

We observe that  $B^{\frac{1}{2}}\rho_N$  converges to  $A_{\varphi}^{1/2}Z$  because Lemma 3 implies that  $B^{\frac{1}{2}}\int \varphi(x-\mu_0) d\{F_{Q,N}(x) - F_Q(x)\}$  is asymptotically normal with mean zero and variance  $A_{\varphi}$ , whereas Lemma 6 implies that

$$B^{\frac{1}{2}}\rho_N = B^{\frac{1}{2}} \int \varphi(x - \mu_0) \, \mathrm{d}\{F_{Q,N}(x) - F_Q(x)\} + o_P(1).$$

We use a Taylor series expansion to write

$$\lambda(T_N) - \lambda(\mu_0) = \lambda'(\mu_0)(T_N - \mu_0) + o_P(|(T_N - \mu_0|).$$

Hence

$$\lambda'(\mu_0)(T_N - \mu_0) + o_P(|T_N - \mu_0|) = B^{-1} \int \varphi(x - \mu_0) \, \mathrm{d}\{F_{Q,N}(x) - F_Q(x)\} + o_P(B^{-\frac{1}{2}}).$$

Taking the absolute value on both sides and using the definition of  $\Psi_N(\mu_0)$  and the fact that  $\lambda(\mu_0) = 0$ , we obtain

$$|T_N - \mu_0| |\lambda'(\mu_0) + o_P(1)| \le |B^{-1} \Psi_N(\mu_0) + o_P(B^{-\frac{1}{2}})|.$$

The right hand side is  $O_P(B^{-\frac{1}{2}})$ , and hence  $(T_N - \mu_0) = O_P(B^{-\frac{1}{2}})$ . Finally

$$B^{\frac{1}{2}}(T_N - \mu_0) = \left[\lambda'(\mu_0)\right]^{-1} B^{\frac{1}{2}} \Psi_N(\mu_0) + o_P(1),$$

completing the proof of Theorem 1.

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